

SPE SoAR

Sean Ford, Bill Walter, and Rob Mellors, LLNL

- Explosion Model Comparison
 - Canonical explosion models compared to SPE
- Moment Tensor Analysis
 - Inversion for SPE moment tensor solution
- Coda-wave Interferometry
 - Inferences on slowness perturbation due to SPE-2/-3
- Spall
 - Standard spall models compared to SPE

An Explosion Model Comparison with Insights from the Source Physics Experiment

Sean Ford and Bill Walter

SSA 2013

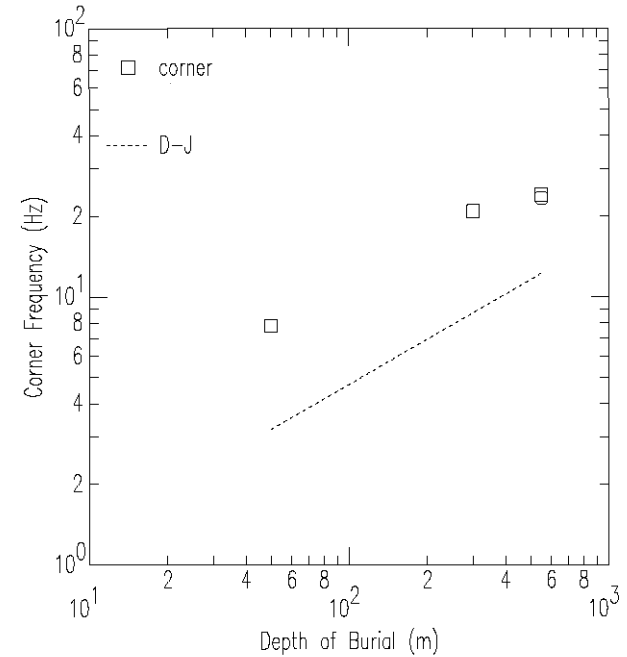
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Motivation & Introduction

- Predict absolute ground motion for explosions using non-standard testing practice (i.e., not NTS or STS media at $\sim 100 \text{ m/kT}^{1/3}$)
- Mueller & Murphy (1971) MM71
 - Scaling laws for seismic observables
- Denny & Johnson (1991) DJ91
 - Regression analysis for seismic observables

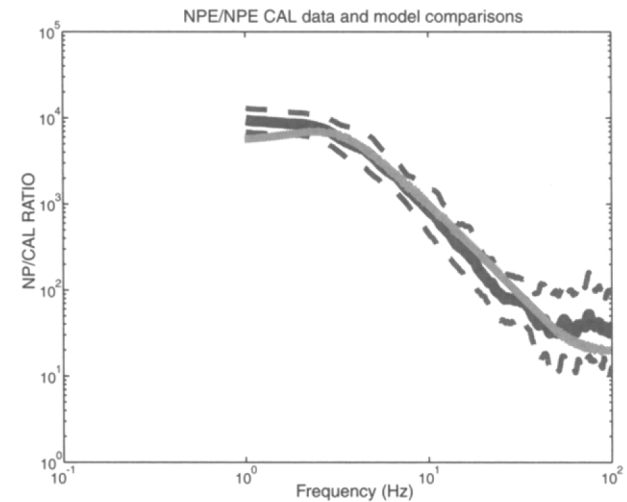
Previous work

- Denny (1998)
 - DOB
 - DJ91 $f_c < \text{observed}$



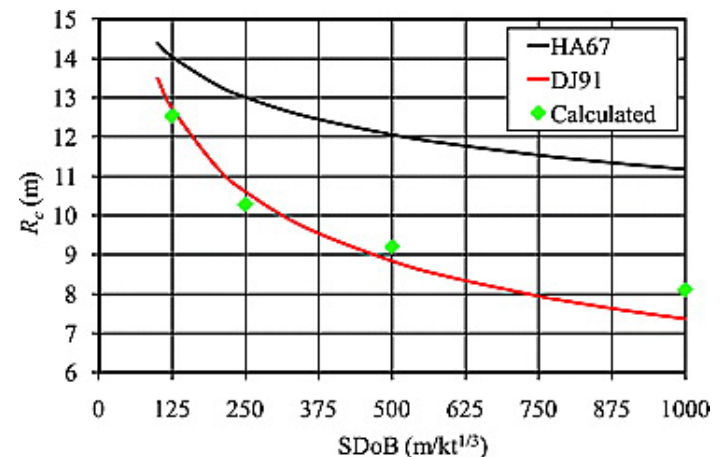
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- Stump et al. (1999)
 - NPE
 - MM71 with $r_c \sim W^{1/3}$



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 - DJ91 $f_c < \text{observed}$
- Stump et al. (1999)
 - NPE
 - MM71 with $r_c \sim W^{1/3}$
- Rougier et al. (2011)
 - Simulation
 - DJ91 $r_c \cong \text{predicted}$



Model comparison (M_0)

- Denny & Johnson (1991) DJ91
 - Combine chem and nuke data in cavity regression where $r_c \sim W^{1/3}$ so $M_0 \sim W$
 - Relate cavity to “measured moment”

$$M_0^{DJ} = 4.2743 \times 10^{10} W \alpha^2 \beta^{-1.1544} P^{-0.4385} 10^{-0.0344 GP} \rho$$

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- Mueller & Murphy (1971) MM71
 - Assume scaling in amplitude and yield for nuke data
 - Employ a cavity regression where
$$r_c \sim W^{0.29} \text{ so } M_0 \sim W^{0.87}$$
$$M_0^{MM} = 3.1416 W^{0.87} \alpha^2 \beta^{-2} R_0^3 P_0 h_0^{1/3} h^{-1/3}$$

Model comparison (f_c)

- DJ91 relate cavity to source radius with $\beta / \pi f_c$

$$f_c^{DJ} = 0.2045 W^{-1/3} \beta^{-0.0642} P^{0.5522} 10^{0.0025 GP} \rho^{-0.7245}$$

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- MM71 relate elastic radius to ω_0 with α / f_c

$$f_c^{MM} = 0.1592 W^{-1/3} \alpha R_0^{-1} h_0^{-1/n} h^{1/n}$$

Model comparison (f_c)

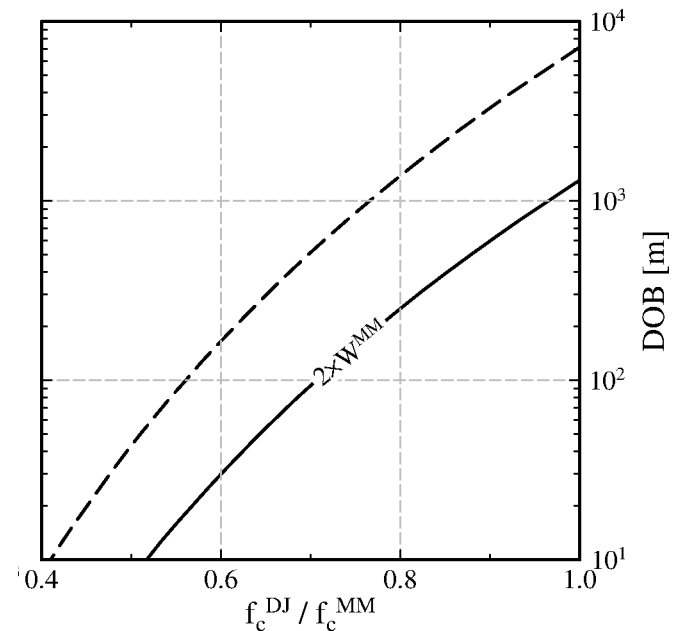
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- DJ91/MM71 $\sim f(\text{DOB})$
 - Allow for chem/nuke factor



Model comparison (W)

- Invert DJ91 and MM71 M_0 relationship for yield

$$W^{DJ} = 2.940 \times 10^{-10} \beta^{1.1544} P^{0.4385} 10^{0.0344 GP} \Psi_{\infty}$$

$$W^{MM} = 4.9207 \rho^{1.1494} \beta^{2.2989} R_0^{-3.4483} P_0^{-1.1494} h_0^{-0.3831} h^{0.3831} \Psi_{\infty}^{1.1494}$$

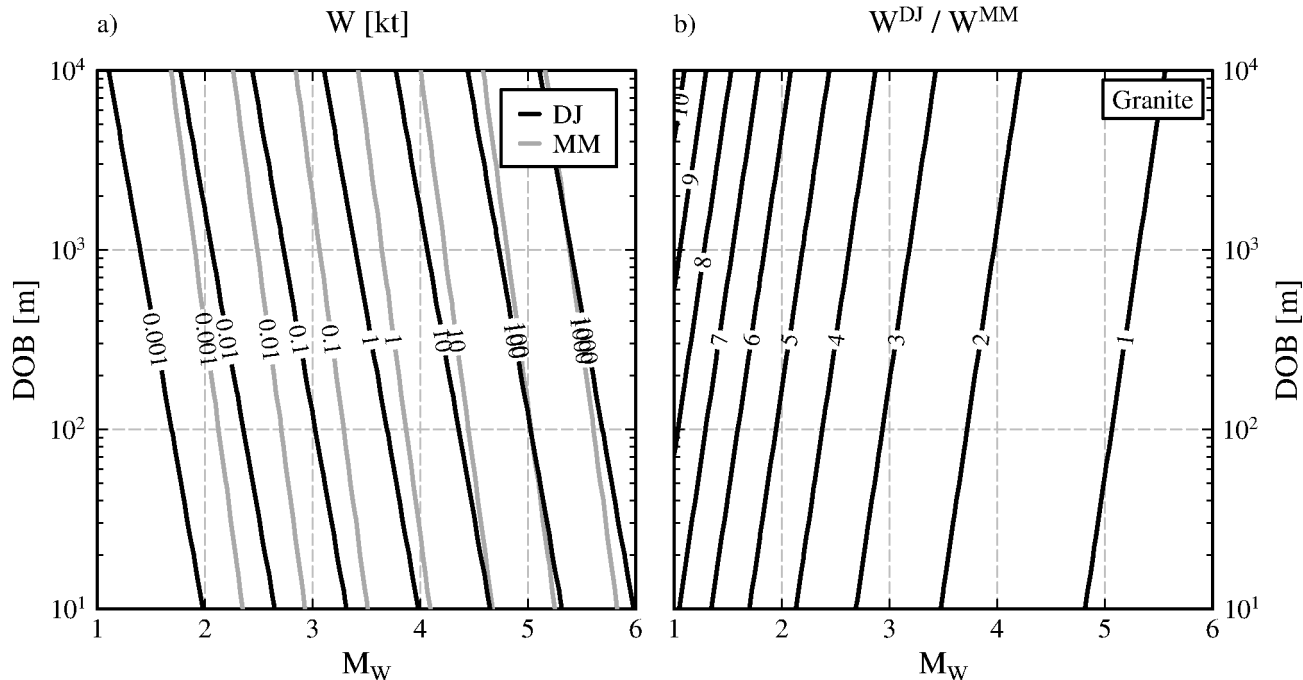
Model comparison (W)

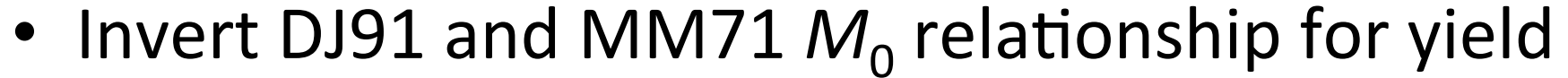
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– Assume shot point $\rho \sim$ overburden ρ $\frac{W^{DJ}}{W^{MM}} = F \frac{h^{0.0554}}{M_0^{0.1494}}$





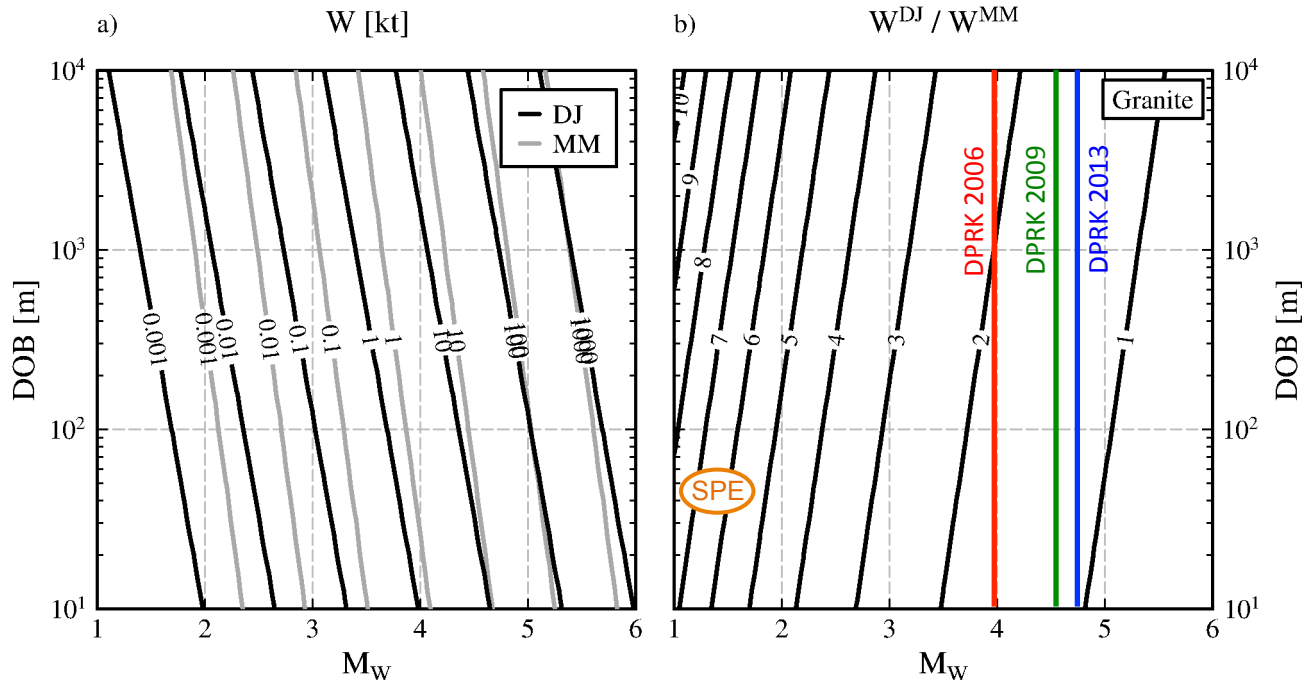
Model comparison (W)

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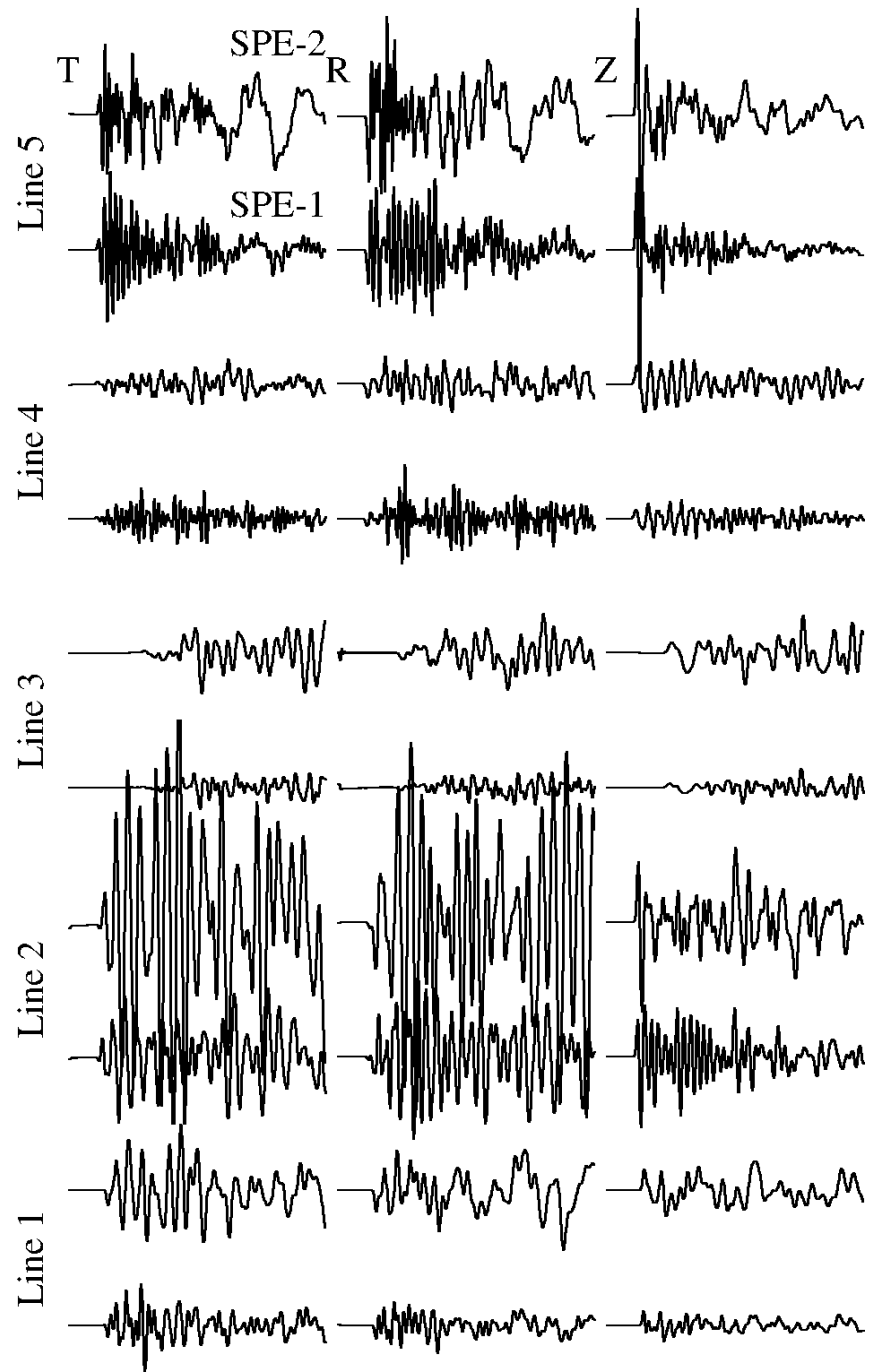
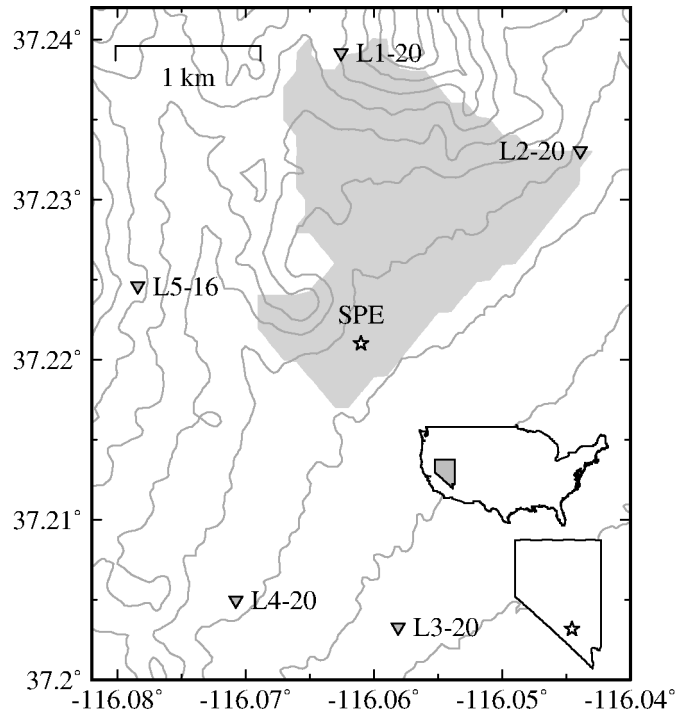
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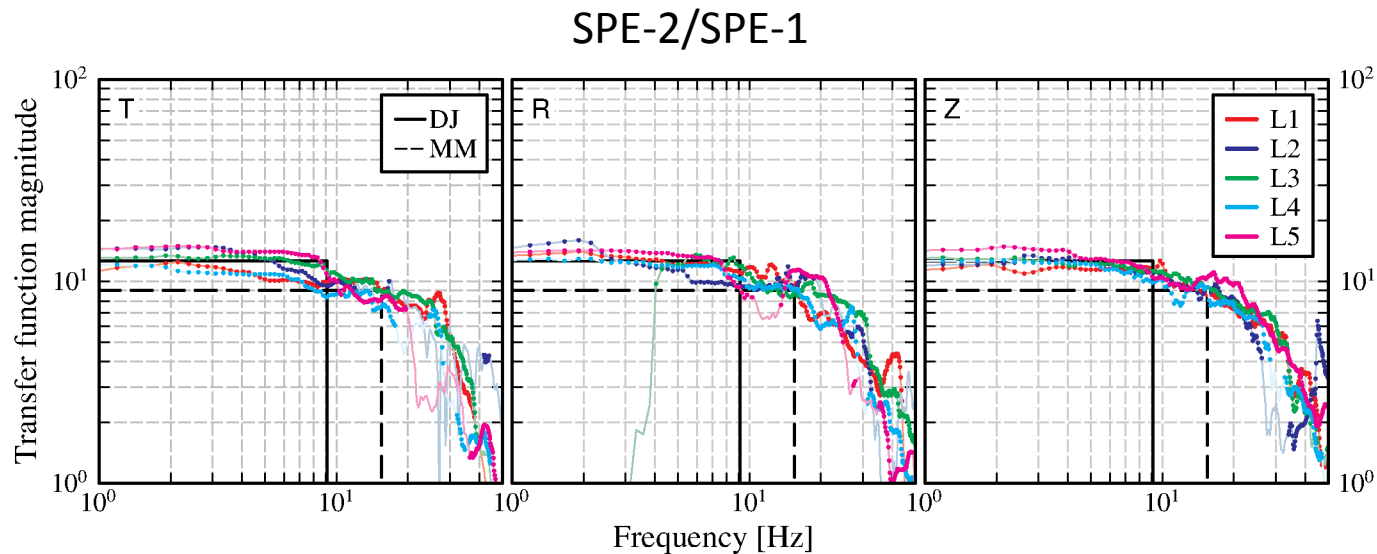


Data: SPE

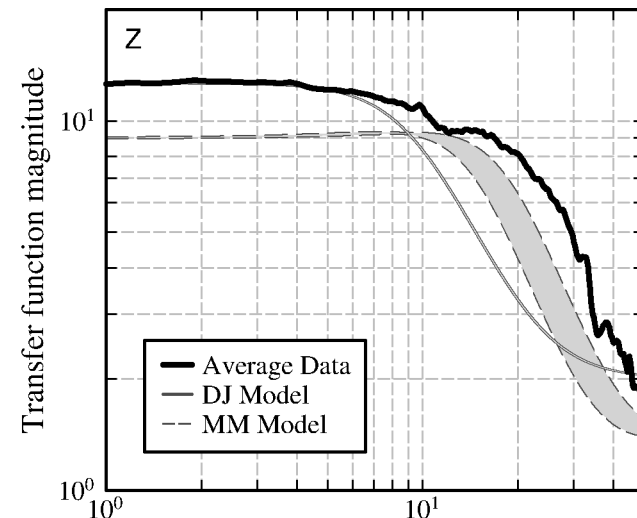


Shot	Yield [kg] (=TNT)	Depth [m]	sDOB [m/kT ^{1/3}]
SPE-1	85.2	54.9	990
SPE-2	991.6	45.7	364

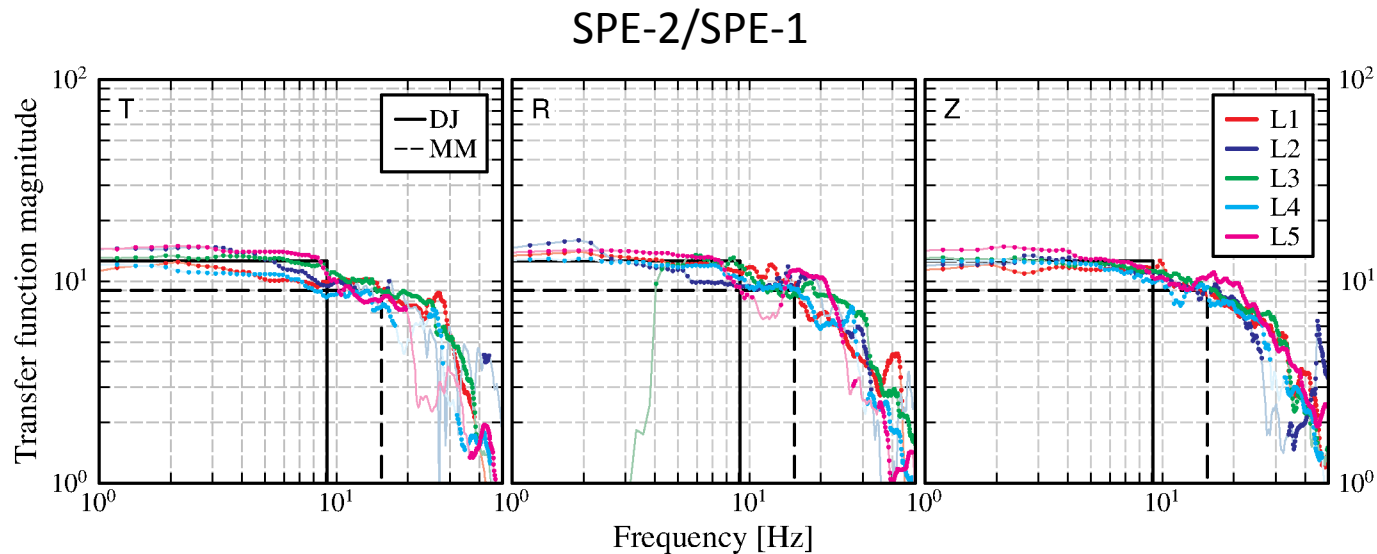
Relative comparison: Spectral ratio



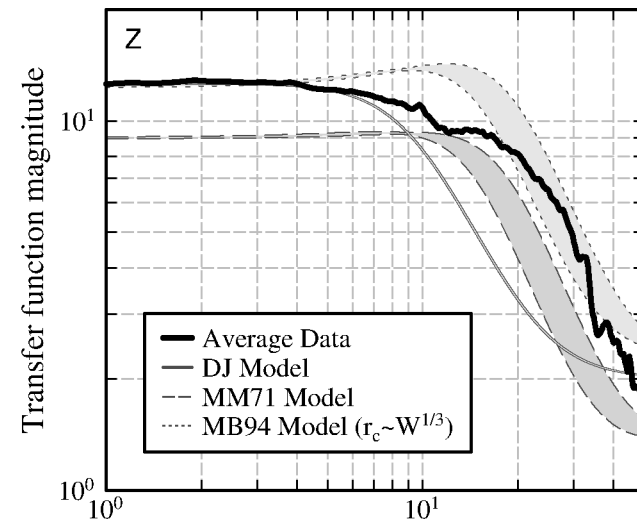
- MM under-predicts long period ratio
- DJ under-predicts corner-frequency



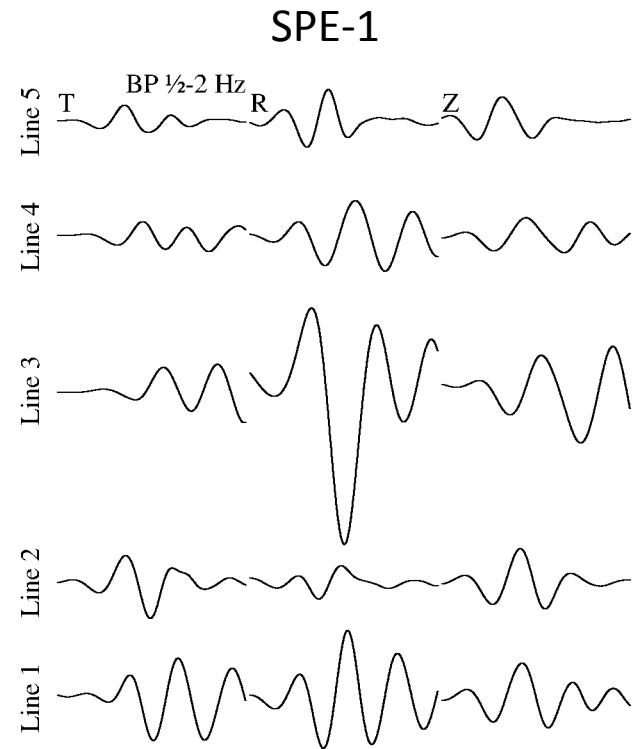
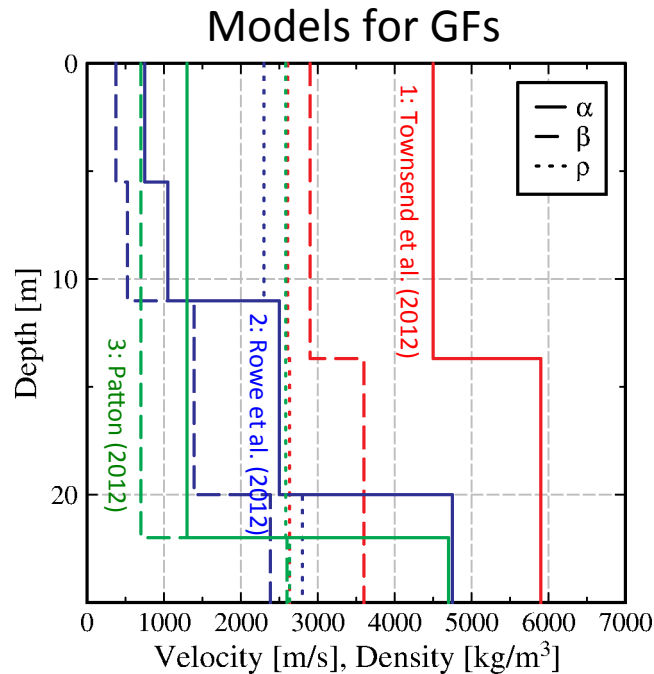
Relative comparison: Spectral ratio



- Altering MM71 so that $r_c \sim W^{1/3}$ (Murphy & Barker, 1994) gives a better fit to the long-period ratio



Absolute comparison: MT synthetic



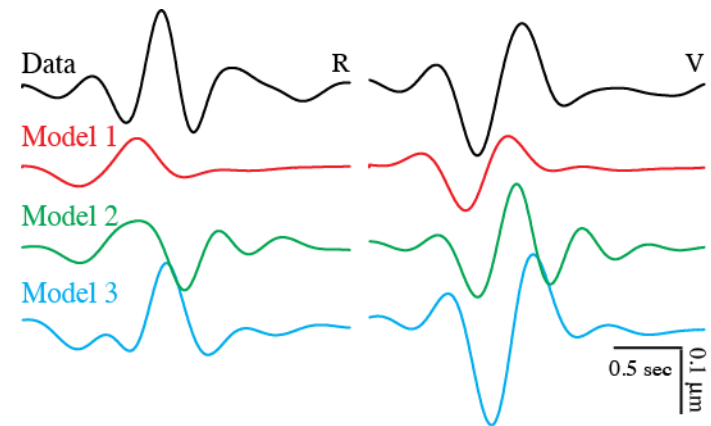
- Energy on tangential component
- Site effects along some lines

Absolute comparison: MT synthetic

SPE-1

Model	M_0^{DJ} [nAk]
1	48.5
2	52.5
3	44.8

Note: 1 Ak = 10^{18} N-m



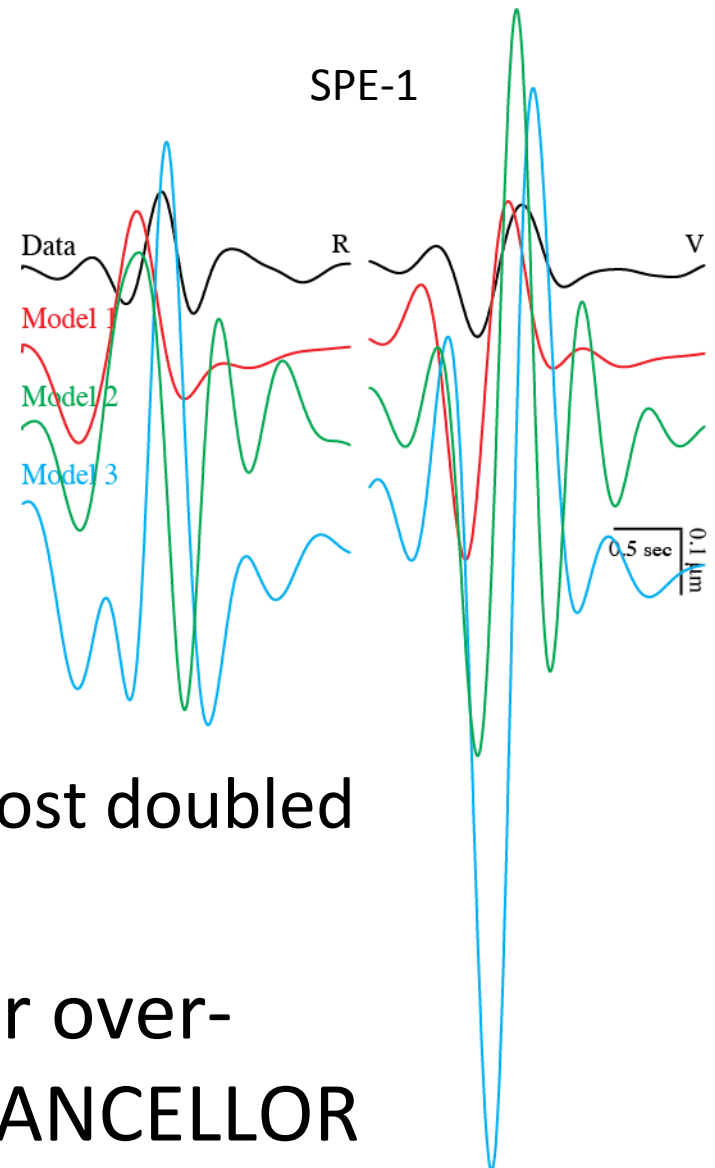
- DJ91 prediction near the observed displacement
- Best phase and amplitude fit by Model 3 (Patton, 2012)

Absolute comparison: MT synthetic

Model	M_0^{DJ} [nAk]	M_0^{MM} [nAk]	
		1×W	2×W
1	48.5	251.7	460.1
2	52.5	373.3	682.3
3	44.8	306.3	559.7

Note: 1 Ak = 10^{18} N-m

- MM71 over-predicts M_0
 - Synthetic amplitudes are almost doubled for chem/nuke factor = 2
- Johnson (1988) found similar over-prediction for HARZER & CHANCELLOR

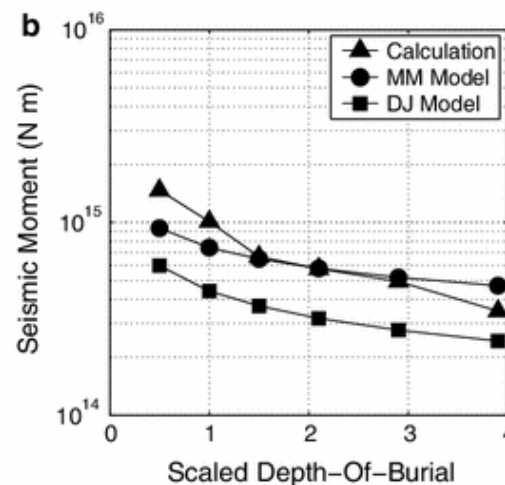
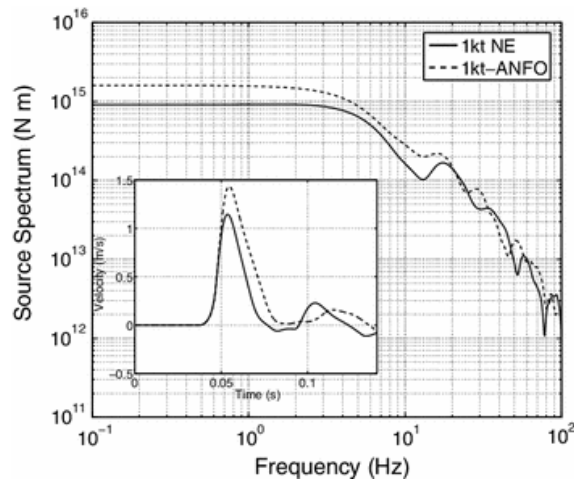


Conclusions & Future work

- Differences in yield predictions are greatest for small and/or deeply buried explosions
- SPE is most consistent with D&J-predicted M_0 and M&M-predicted f_c

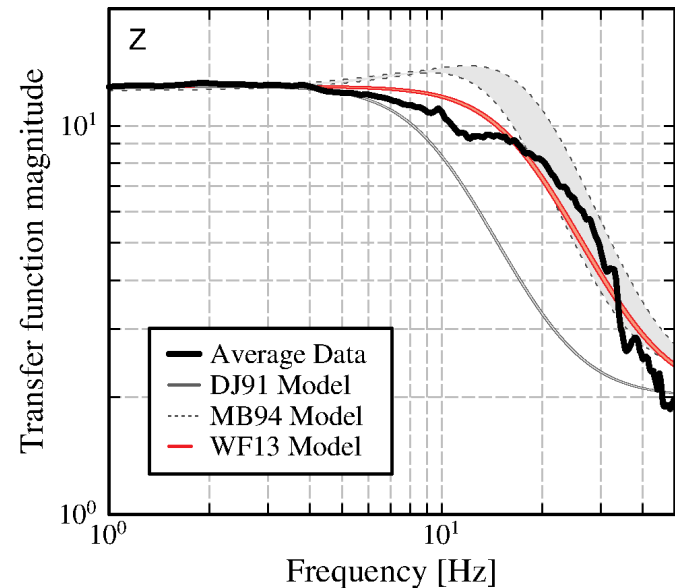
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- Future work: Chem/Nuke at depth (Xu et al., '12)



Conclusions & Future work

- Differences in yield predictions are greatest for small and/or deeply buried explosions
- SPE is most consistent with D&J-predicted M_0 and M&M-predicted f_c
- Future work: New model



Moment tensor analysis of SPE-1 and -2

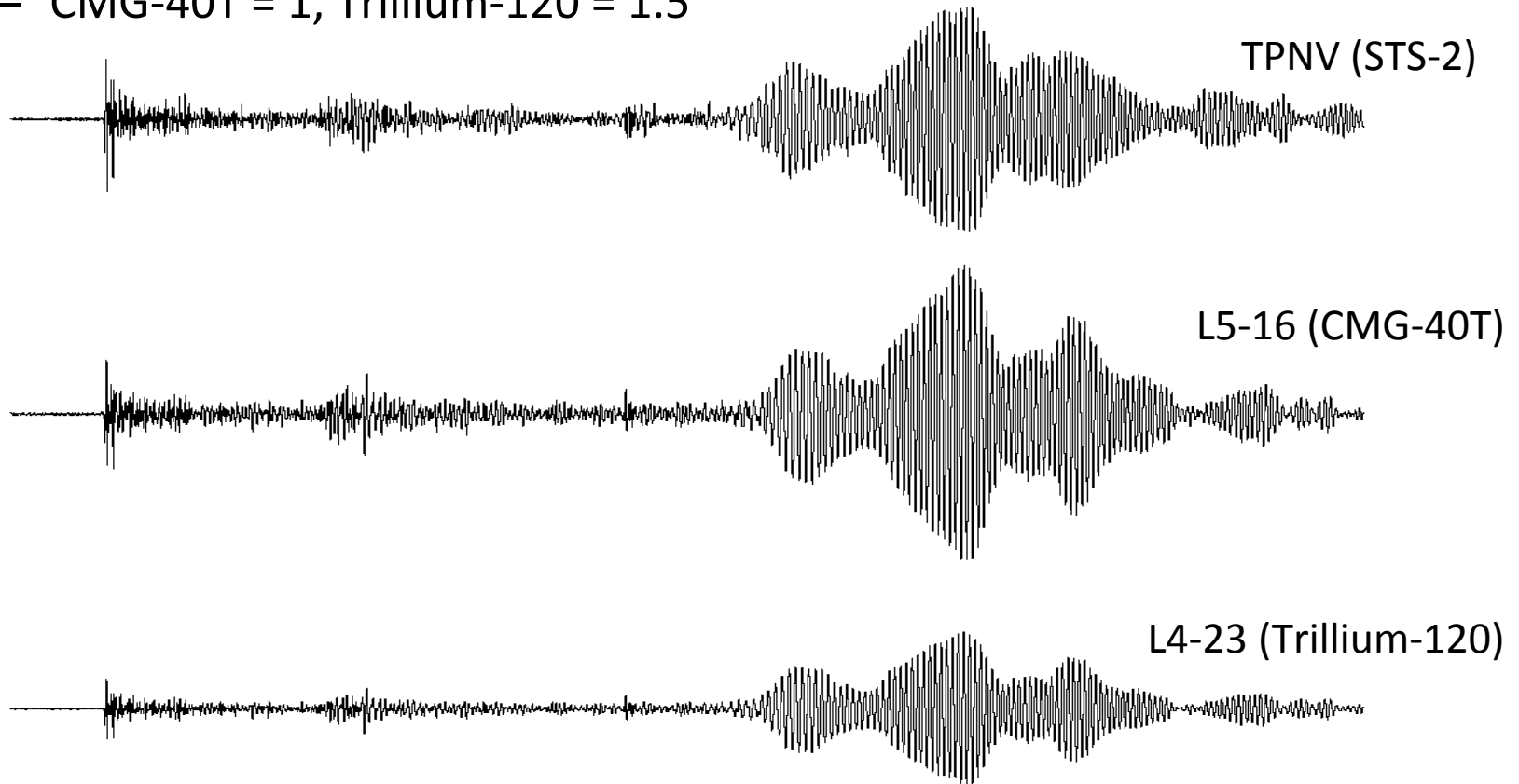
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SSA 2012

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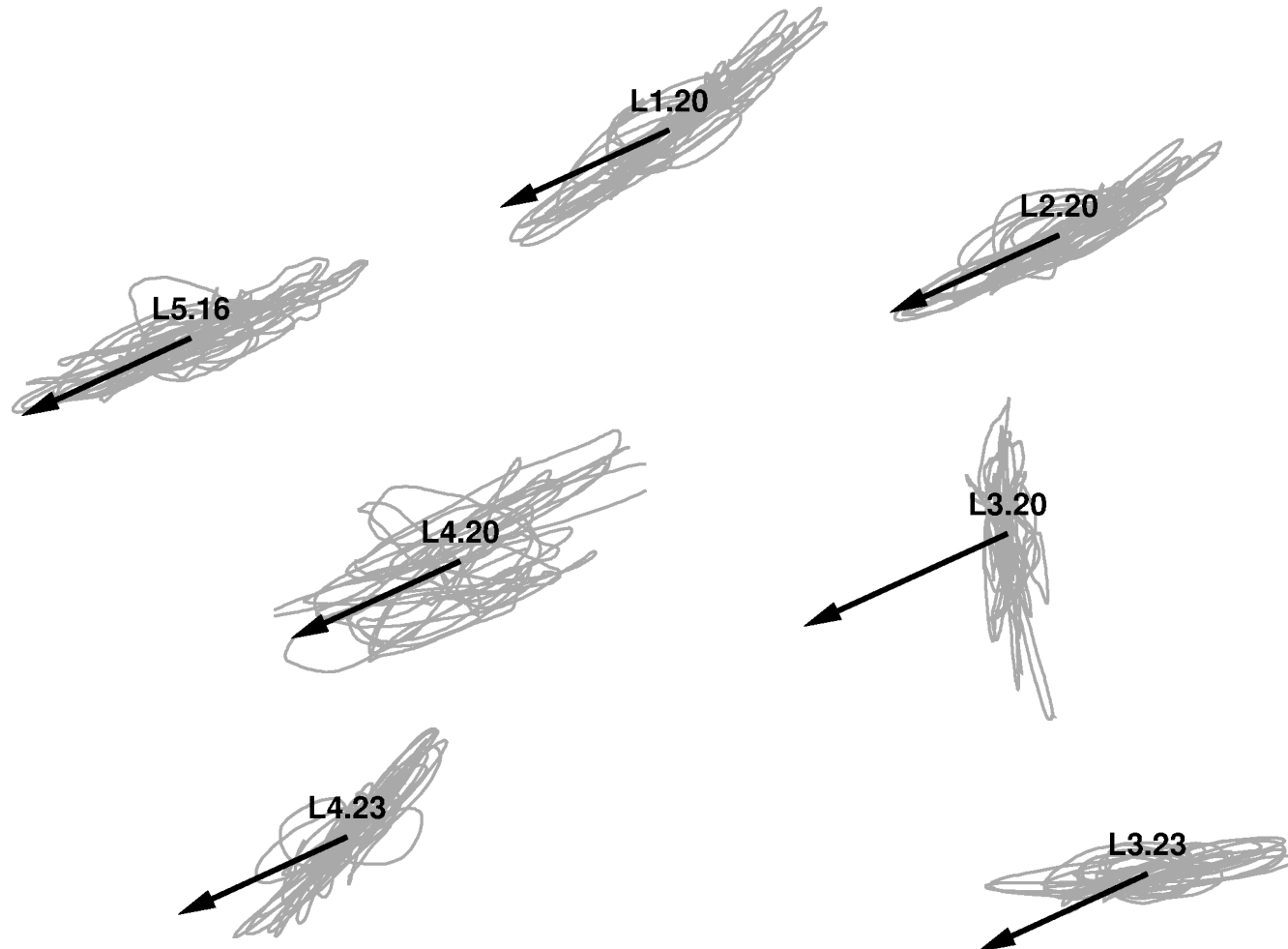
Calibration 1: Response

- H/V ratios of regional events similar between SN and TPNV
 - H/V noise ratio SN=5 TPNV=1¼ (much noisier horizontals)
- Long-period (15-25s) ratio of SN to TPNV teleseismic event provided ground-motion gain corrections
 - CMG-40T = 1, Trillium-120 = 1.5



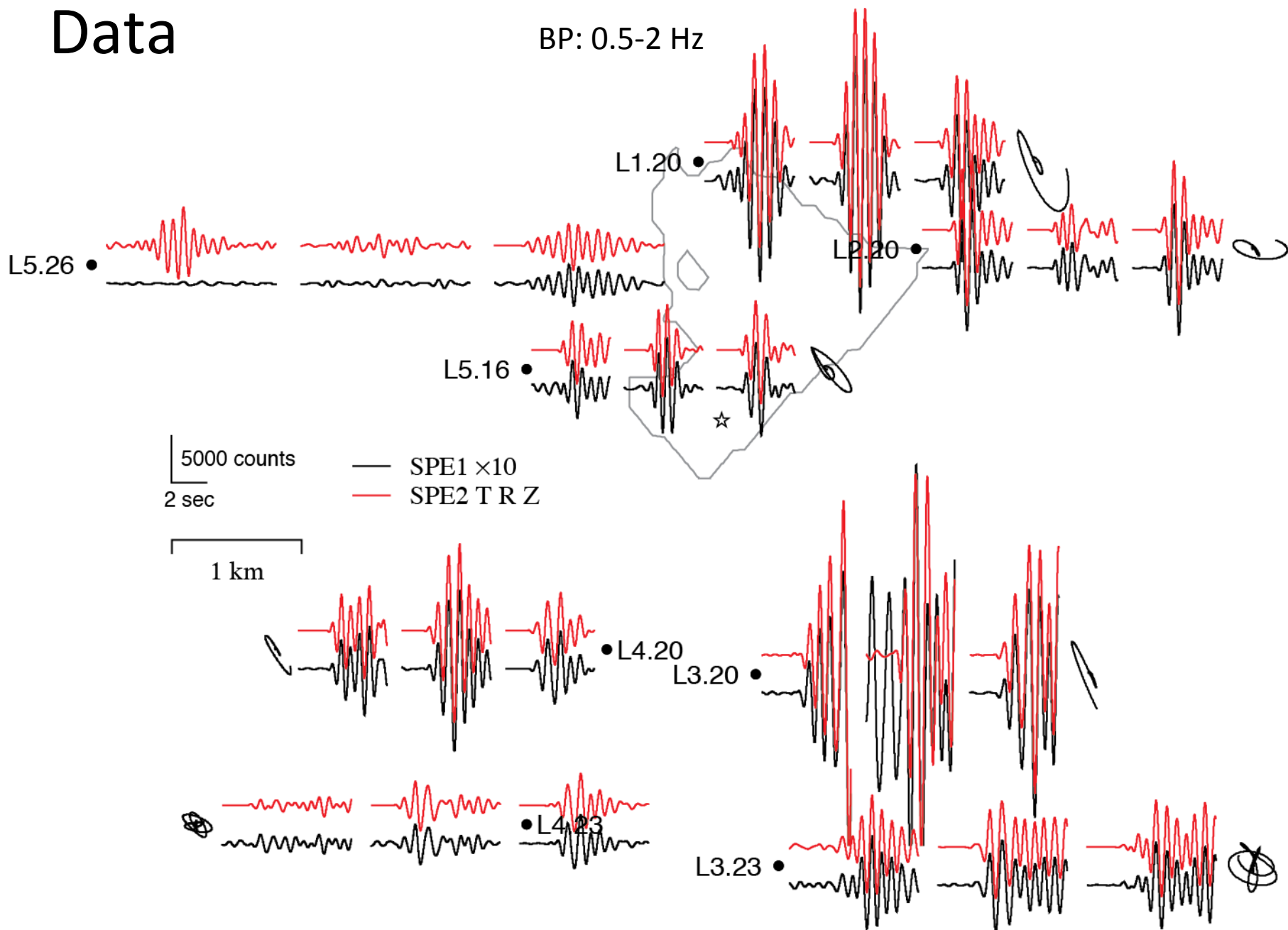
Calibration 2: Orientation

- Horizontal particle motion of teleseismic event
 - $L1-20/L2-20/L3-20/L4-20/L5-16 = +5/0/?/0/-1$



Data

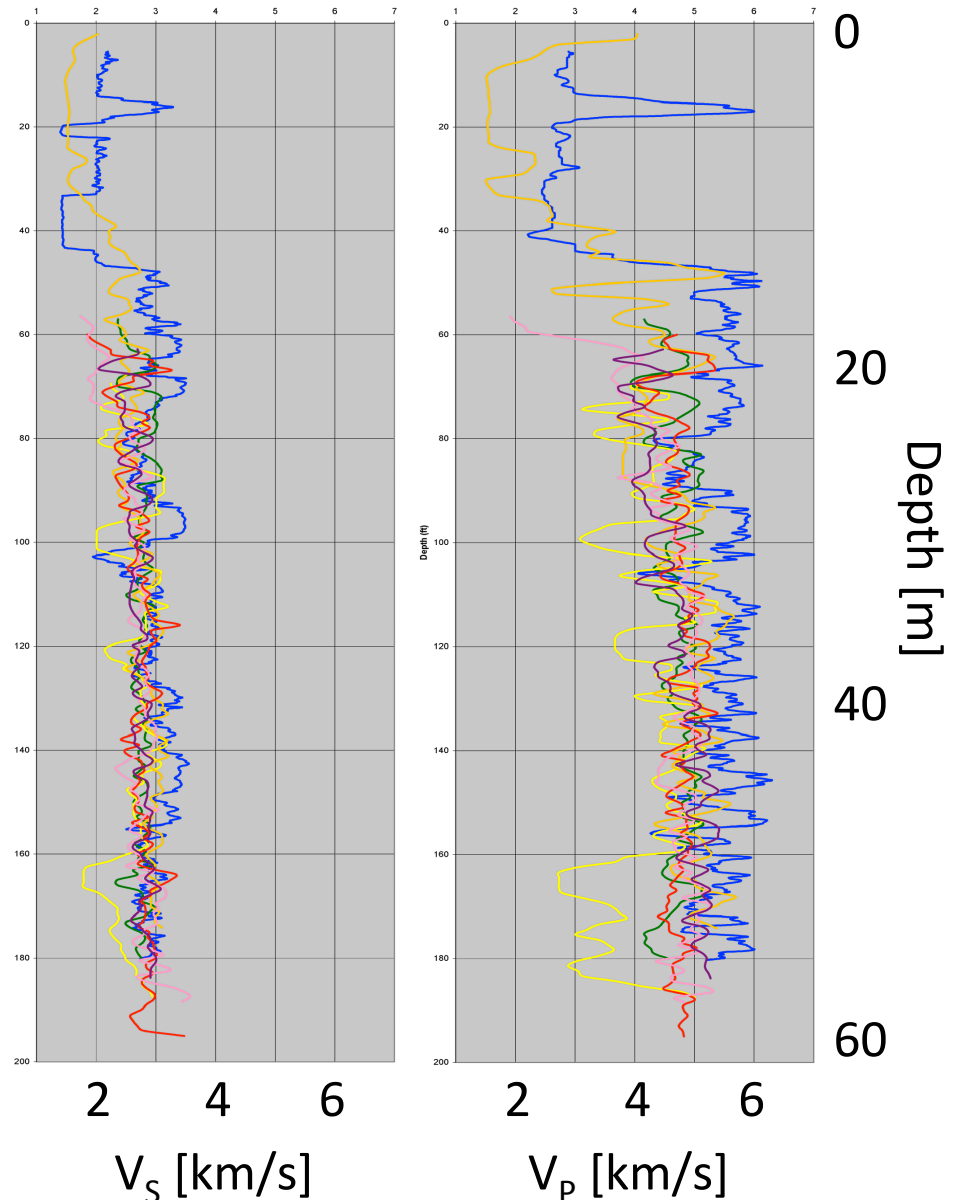
BP: 0.5-2 Hz



Velocity model

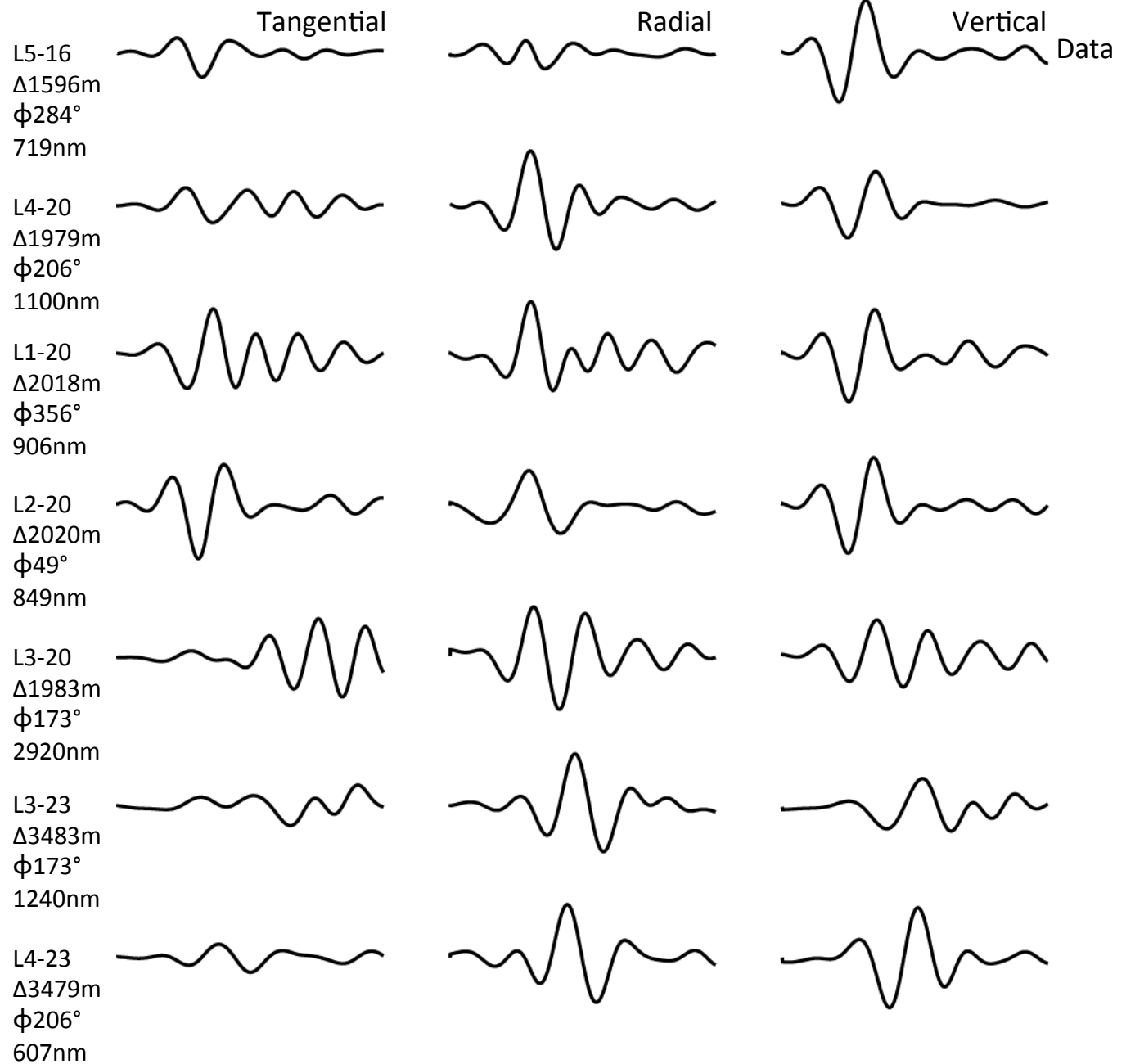
- $V_p/V_s = 5000/2700$ [m/s]
 - $V_p/V_s = 1.852$ $v = 0.294$
- 20m thick weathered layer
 - $V_p/V_s = 2900/1600$ [m/s]
- $\rho = 2600/2650$ kg/m³
- $Q = 600 / 300$
- Constant to depth

Well log (Townsend et al., 2012)



Data

3-component
Displacement
0.5 – 2 Hz

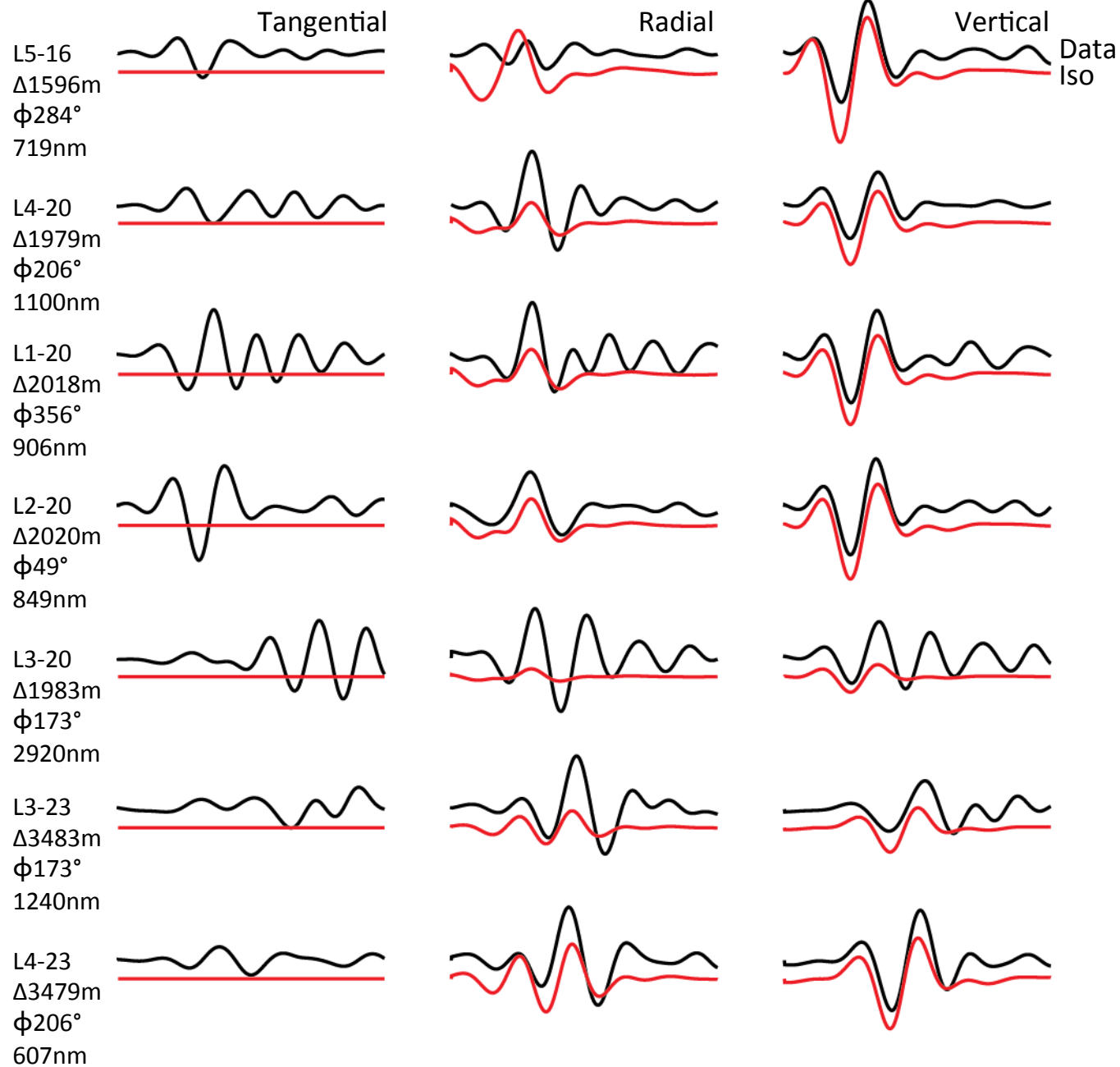
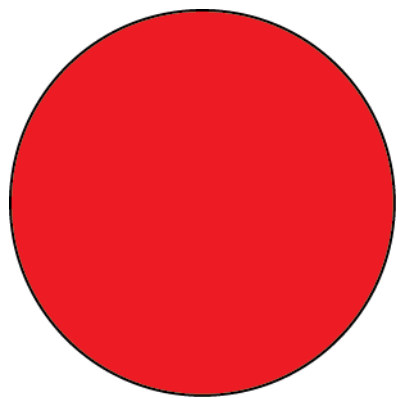


Explosion

$M_0 6.3 \times 10^{11}$ N-m

$M_W 1.8$

VR 44.5%



— 1 sec

Full MT

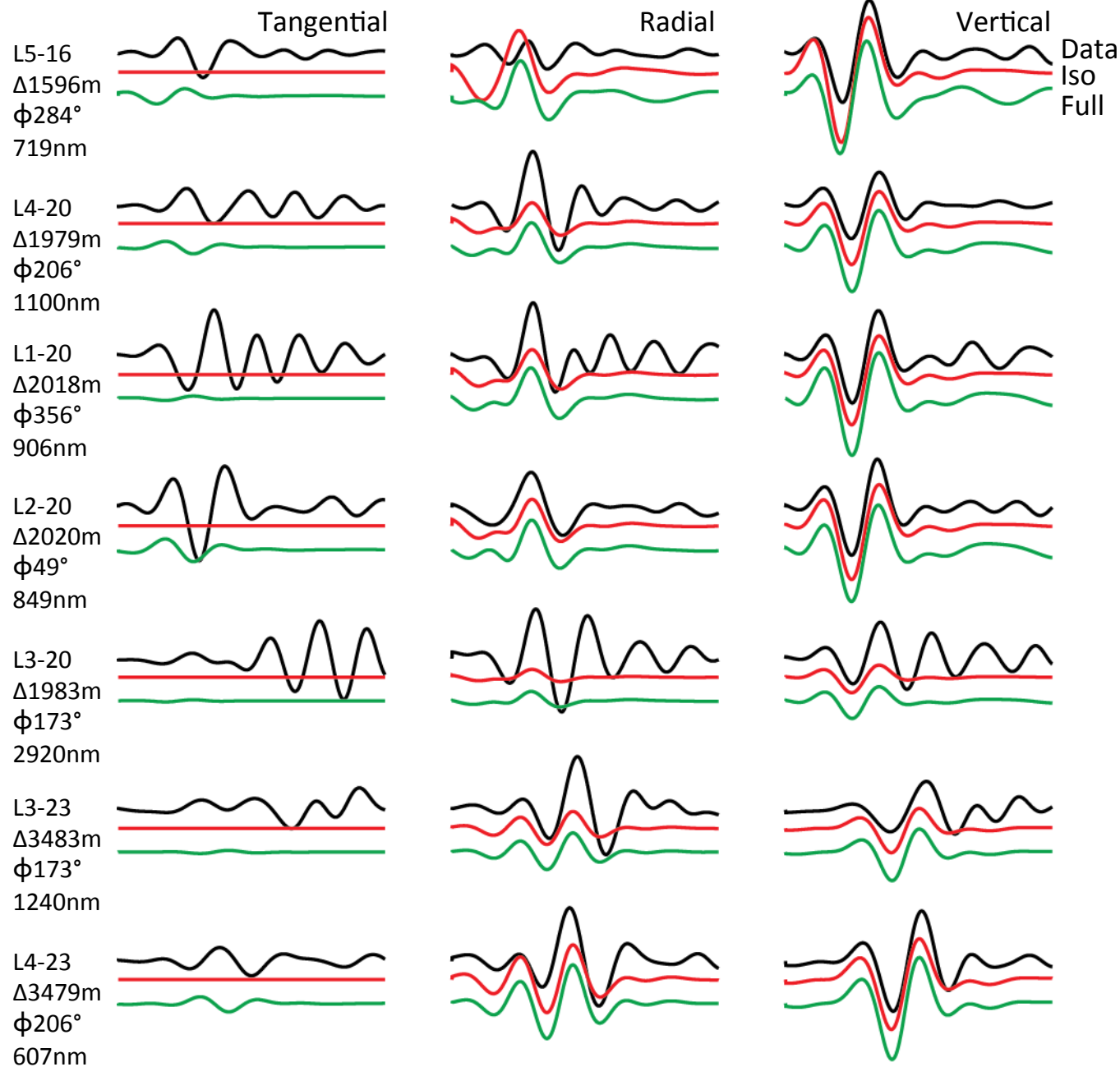
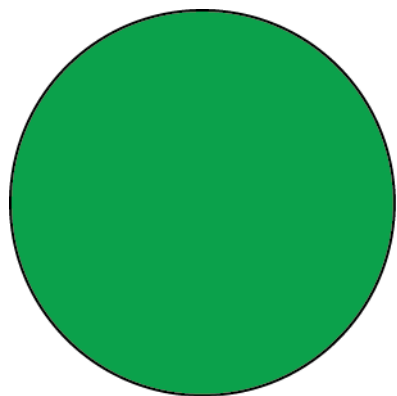
$M_0 2.1 \times 10^{12}$ N-m

$M_W 2.2$

VR 48.4%

ISO/CLVD/DC

75 / 20 / 5



Deviatoric

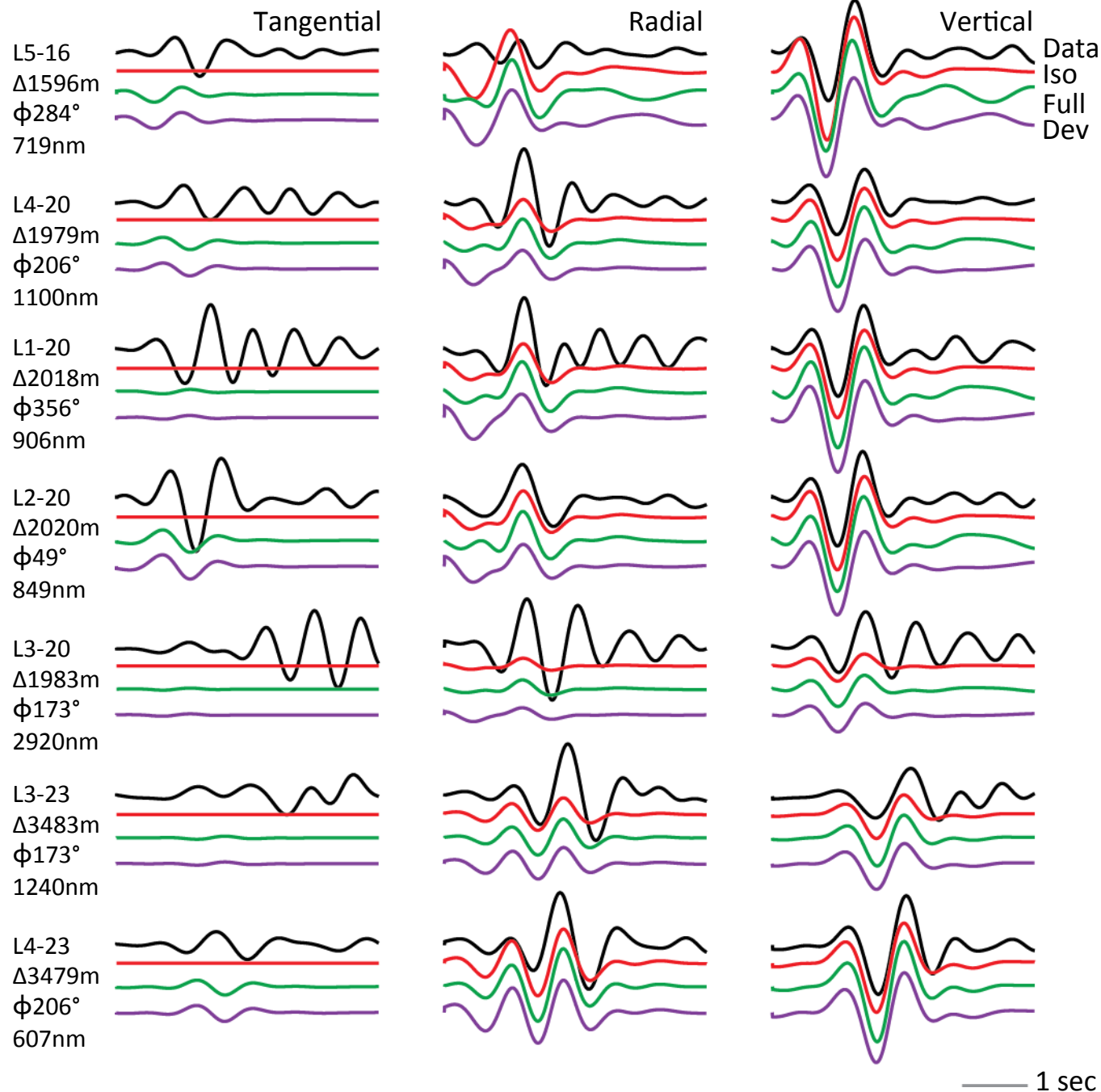
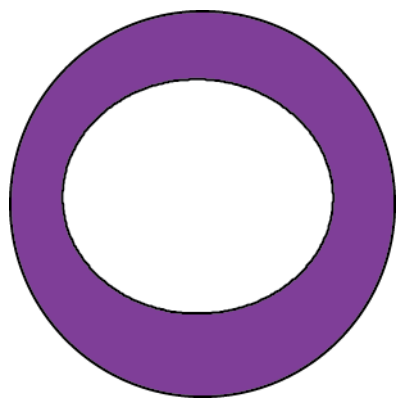
$M_0 4.1 \times 10^{11}$ N-m

$M_W 1.7$

VR 42.3%

CLVD/DC

68 / 32

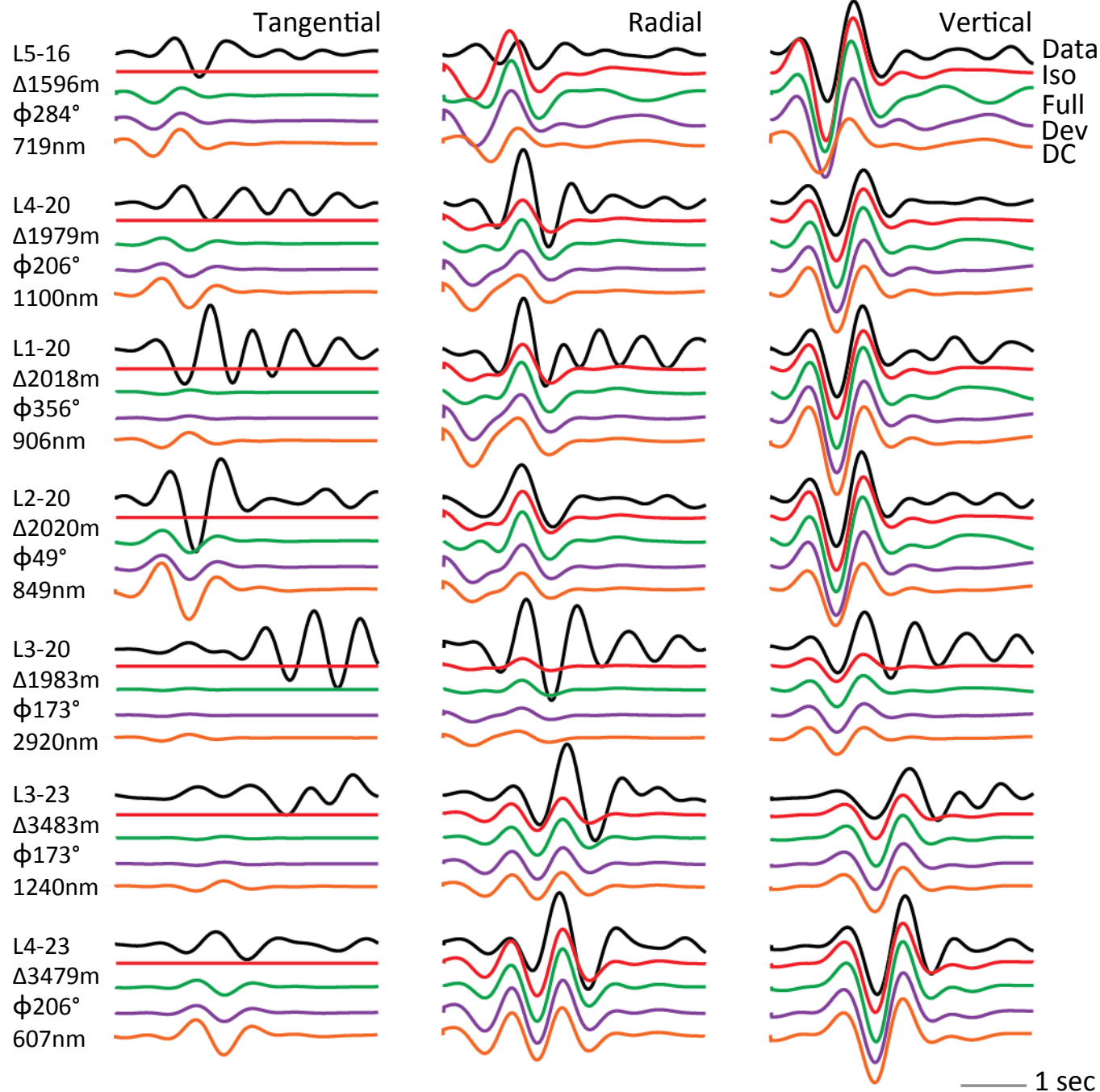
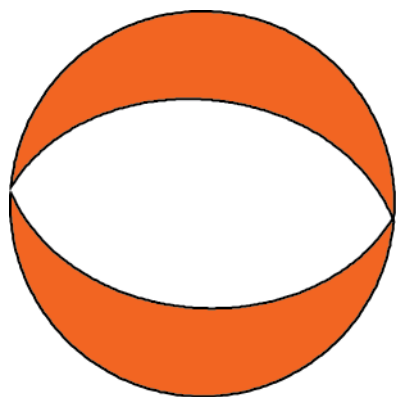


Best-DC

$M_0 3.1 \times 10^{11}$ N-m

$M_W 1.6$

VR 36.7%



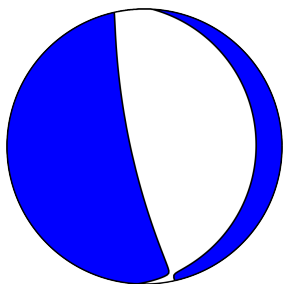
Explosion

$M_0 10.9 \times 10^{11}$ N-m

$M_W 1.9$

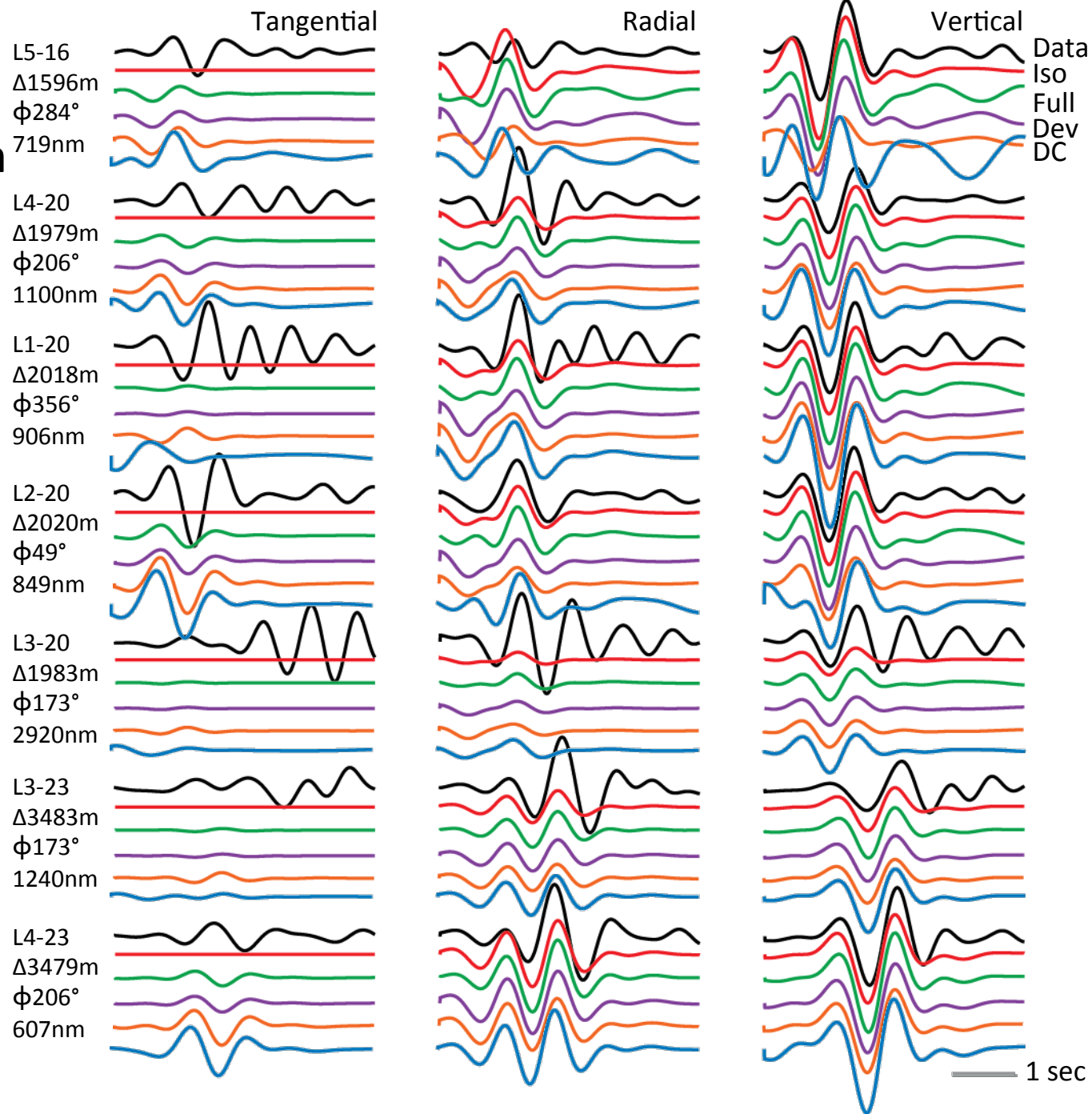
60%

+ DC

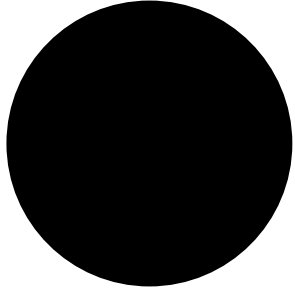


$M_0 8.6 \times 10^{11}$ N-m

$M_W 1.7$



Explosion

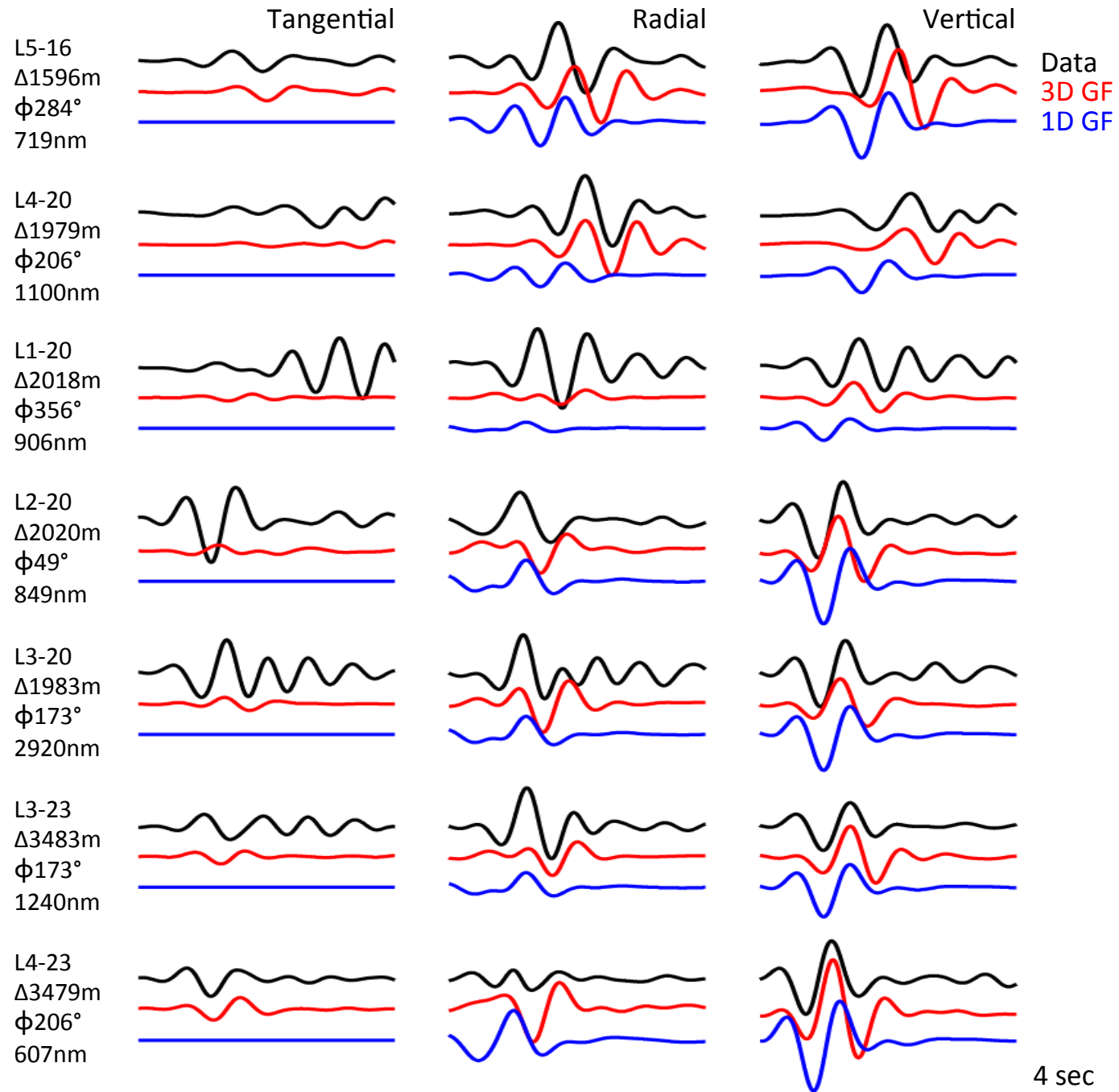


1D GFs

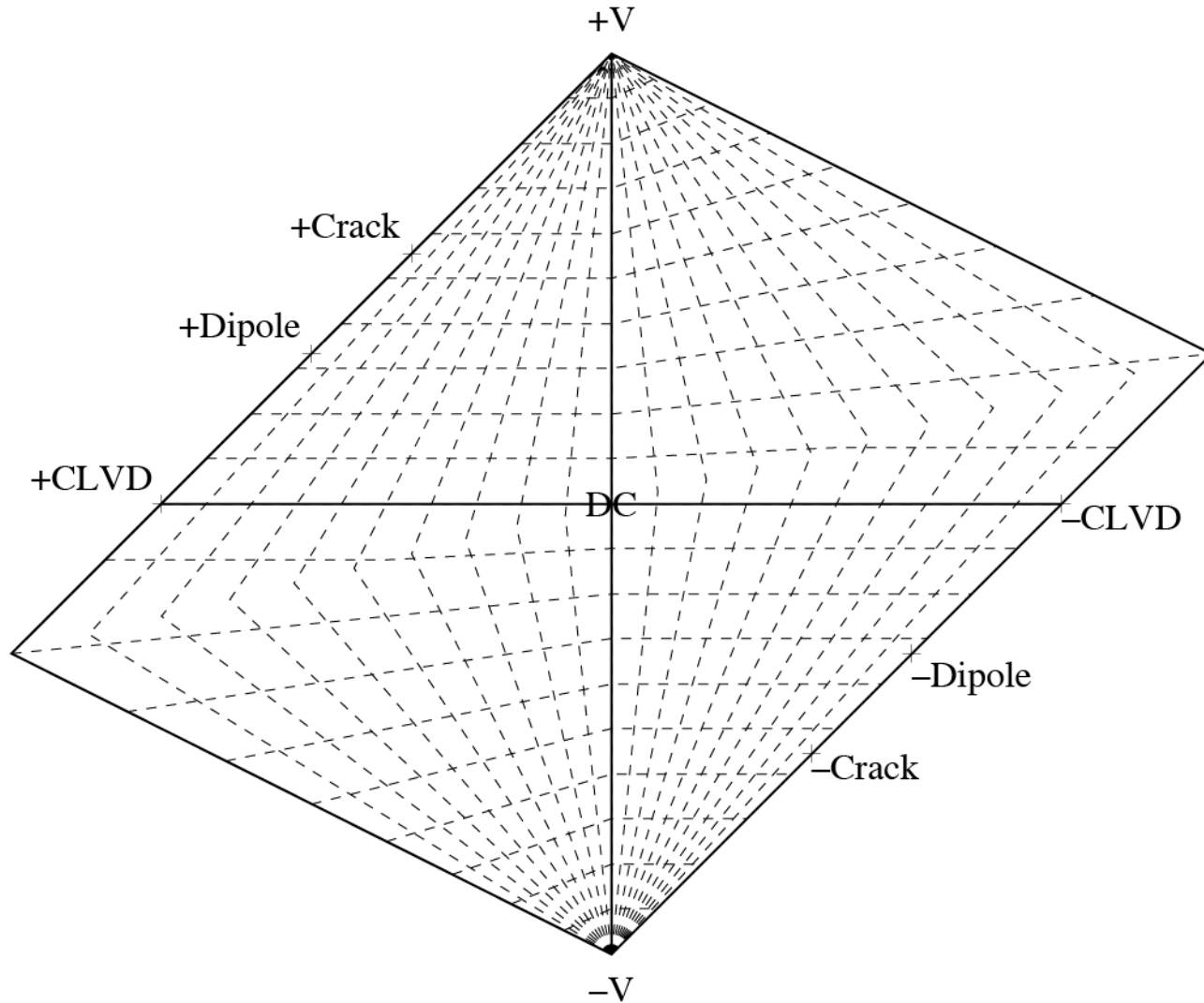
$M_0 63 \times 10^{10} \text{ N-m}$

3D GFs

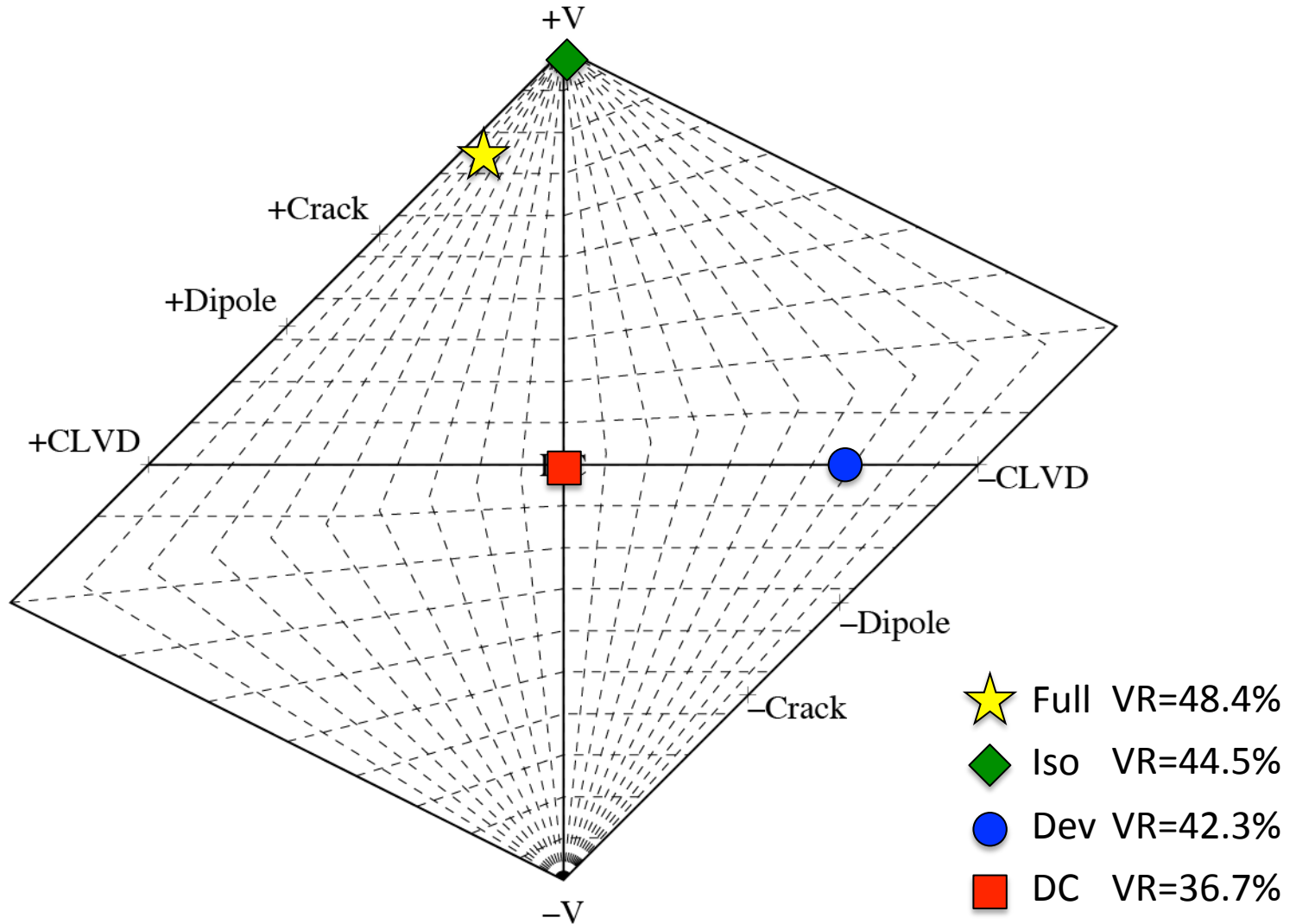
$M_0 38 \times 10^{10} \text{ N-m}$



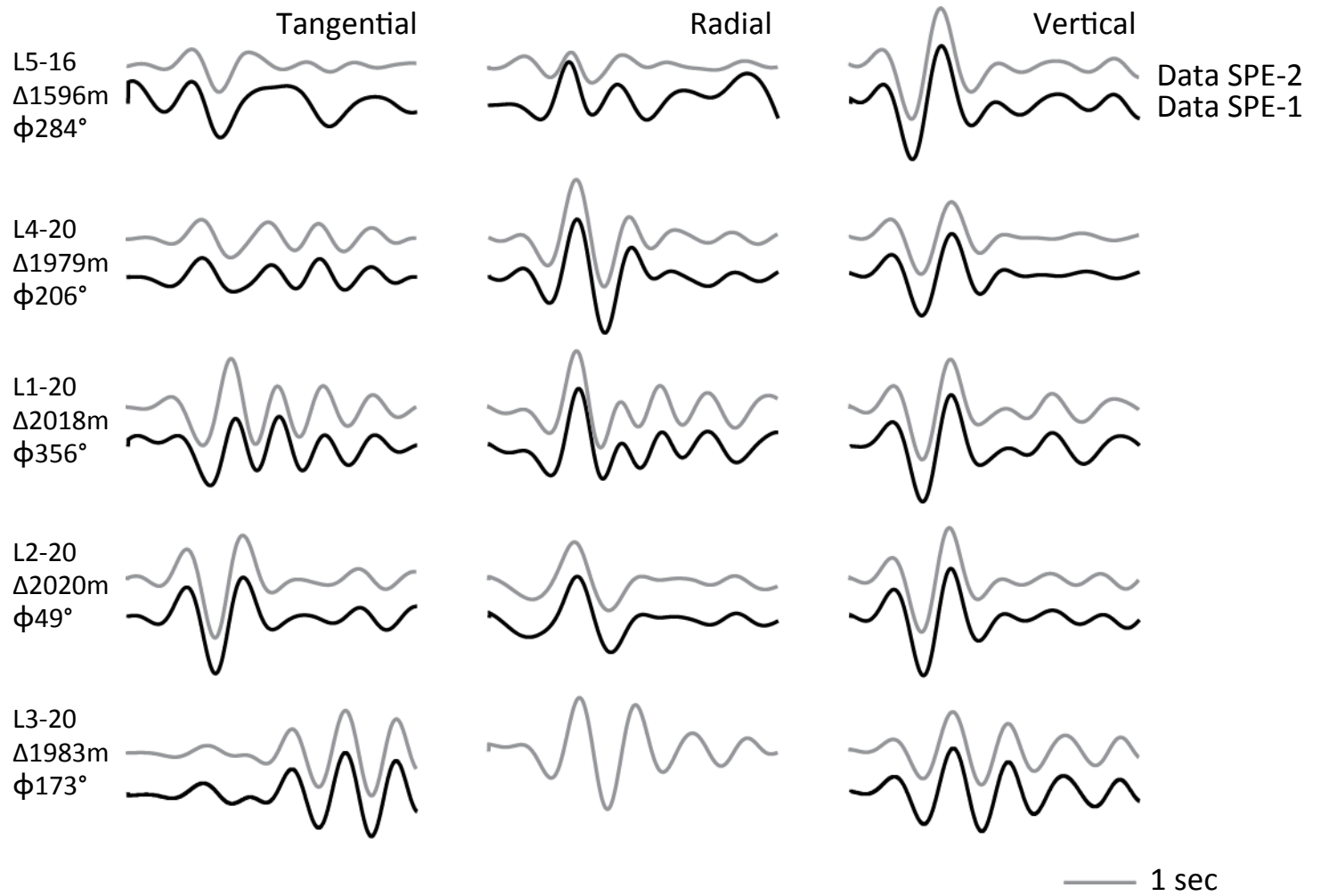
Source-type plot (Hudson et al., 1989)



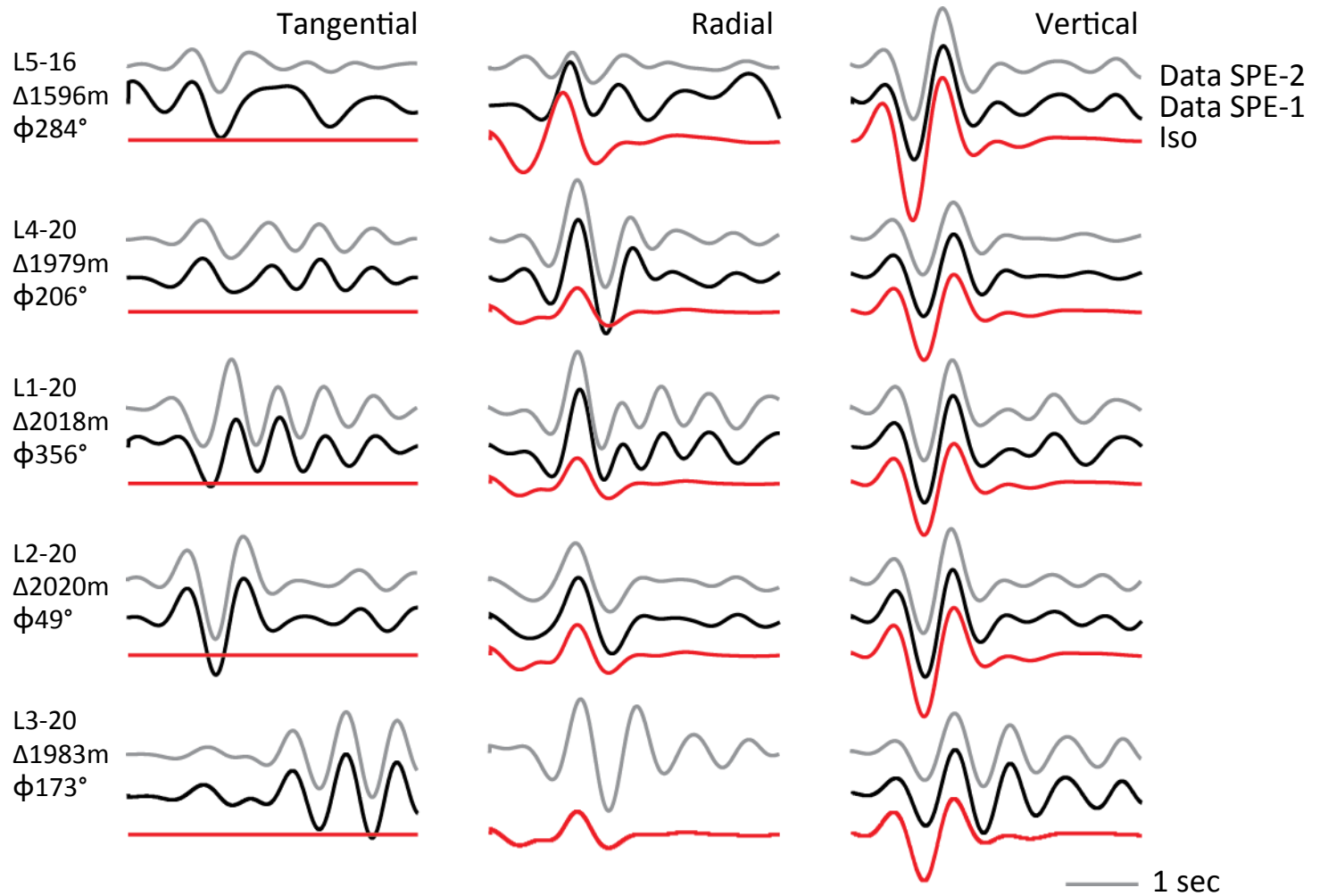
Source-type plot (Hudson et al., 1989)



SPE-1



SPE-1 Explosion - 6×10^{10} N-m ($M_W 1.2$)



- $M_0(\text{SPE-2}) \cong 10 \times M_0(\text{SPE-1})$

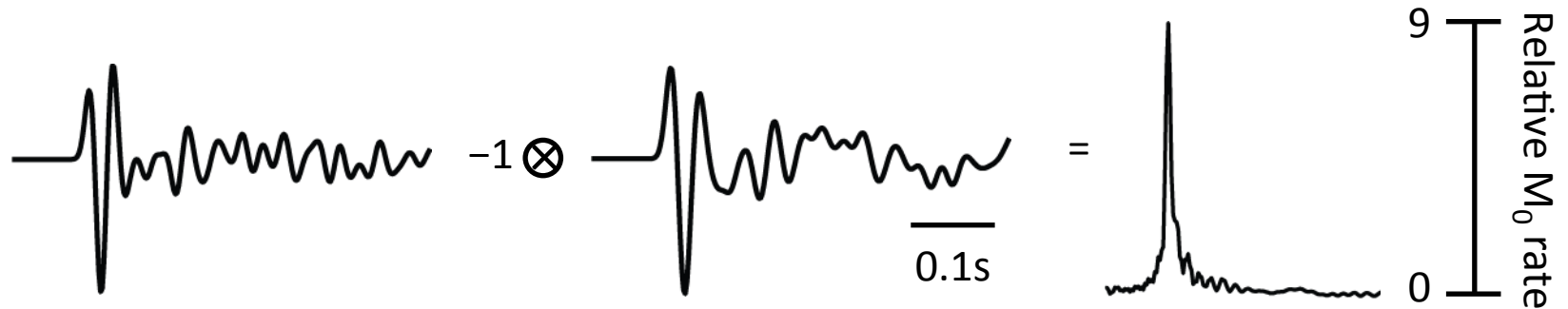
Deconvolution for relative moment-rate

- Deconvolve SPE-1 from SPE-2 on L5-16 vertical at <40 Hz

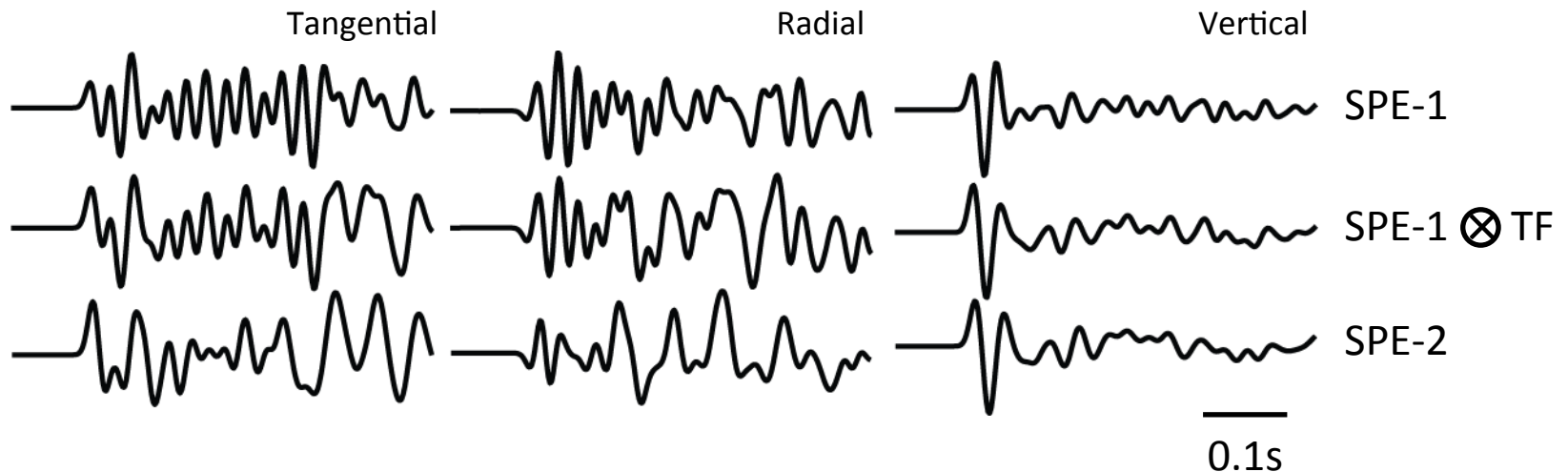


Deconvolution for relative moment-rate

- Deconvolve SPE-1 from SPE-2 on L5-16 vertical at <40 Hz



- Convolve moment-rate function with SPE-1 L5-16 all channels



Denny & Johnson (1991) scaling relations

- Regression for cavity radius as a function of yield obtained

$$R_c = \frac{1.47 \times 10^4 W^{1/3}}{\beta^{0.3848} P_0^{0.2625} 10^{0.0025 GP}}$$

- Regression for seismic moment obtained

$$M_0 = \frac{1}{311} M_t P_0^{0.349} 10^{-0.027 GP} \quad \text{where} \quad M_t = \frac{4}{3} \pi \rho \alpha^2 R_c^3$$

- Prediction of R_c for SPE-1/-2 = 0.79 / 1.79 m
- Explosion M_0 predicts W for SPE-1/2 = 98 / 945 kg TNT

Preliminary findings

- SPE-2 best-fit isotropic M_0 6.3×10^{11} N-m = $10 \times$ SPE-1 isotropic M_0
- SPE-1/-2 low-frequency waveforms are very similar
- Full MT low-frequency point-source inversion cannot fit tangential displacement
- Evidence for multipathing requires >1D velocity-model

Future work

- Distributed source
- Variable 1D models to capture M_0 error
- Higher frequency and greater bandwidth inversions
- Variable weighting of low-amplitude initial arrivals
- Use of eGF for kinematic source inversion

Coda wave interferometry of SPE-2/-3 for velocity change due to damage

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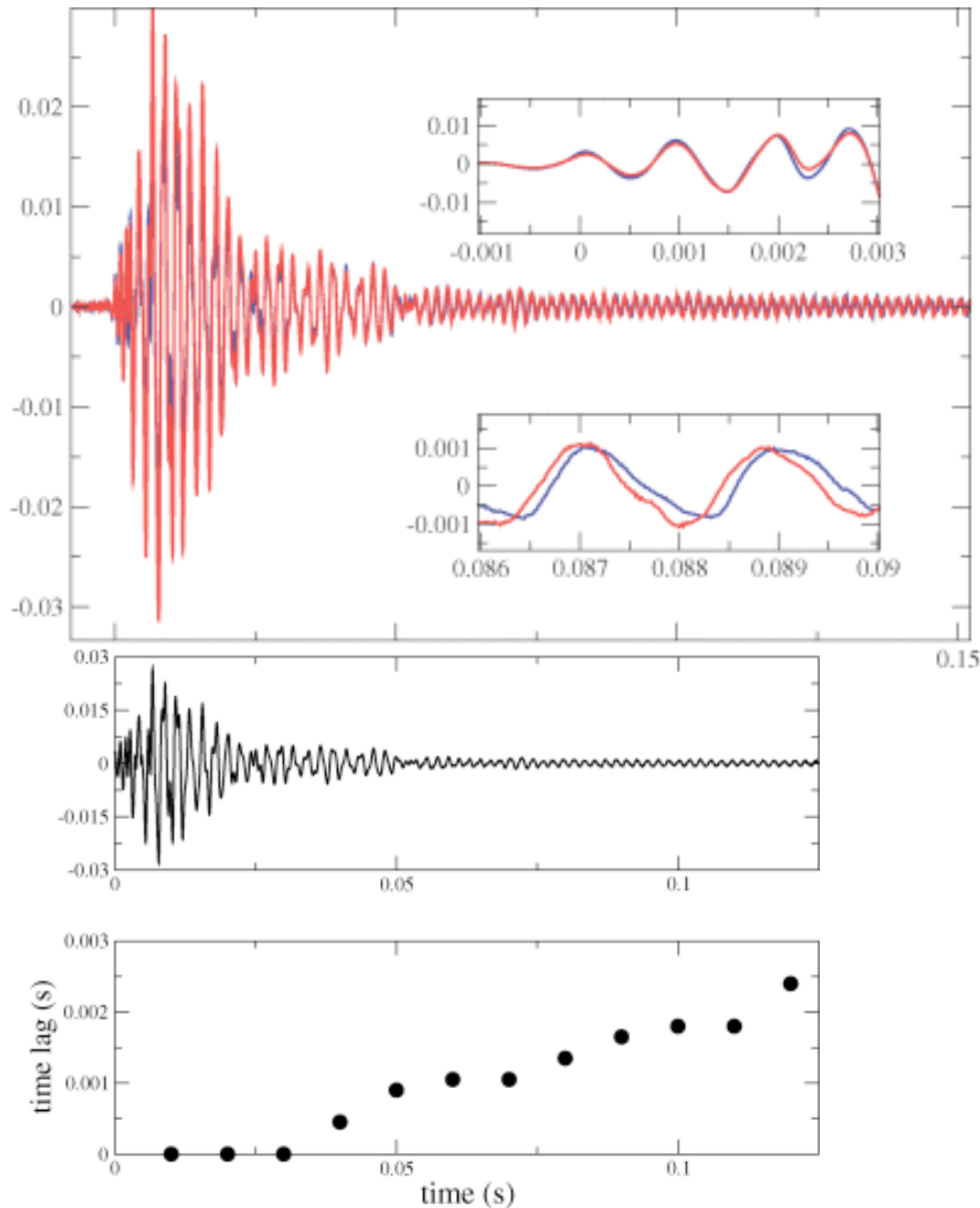
Approach

- GS11D-vertical observations of SPE-2/-3
- Non-overlapping windows of 0.4 s ($10 \times T_{\text{dom}}$)
- Damage should change elastic properties leading to velocity change and travel-time perturbation
- Use coda wave interferometry to measure travel-time perturbation [Snieder et al., 2002]

$$\frac{\delta v}{v} = -\frac{\delta t}{t}$$

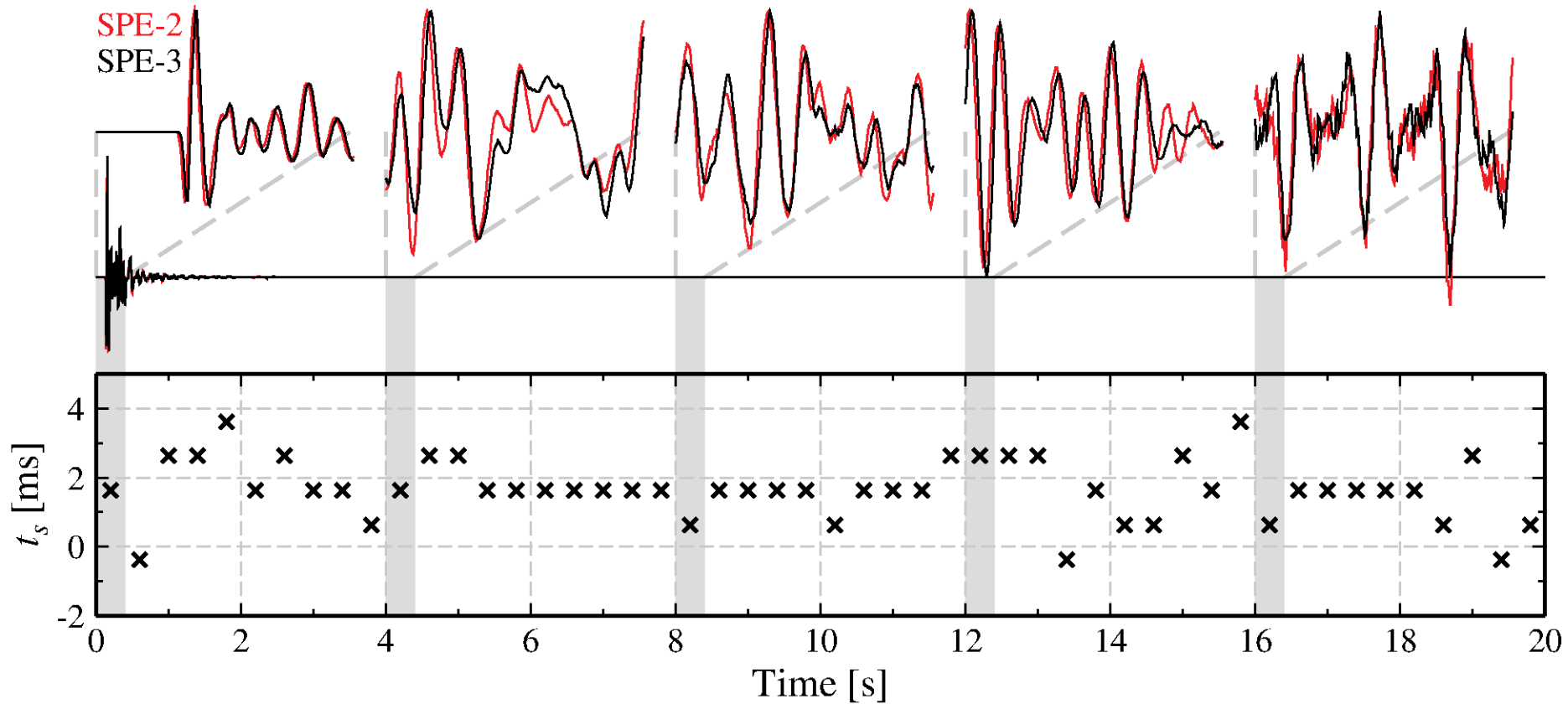
**Example from Gret et al.
[2006] *Geophys. J. Int.***

Two waveforms, one measured at 4.14 MPa of pressure in the cell (blue) and one measured at 12.41 MPa (red). The upper panel shows the early time window and the lower panel part of the coda.

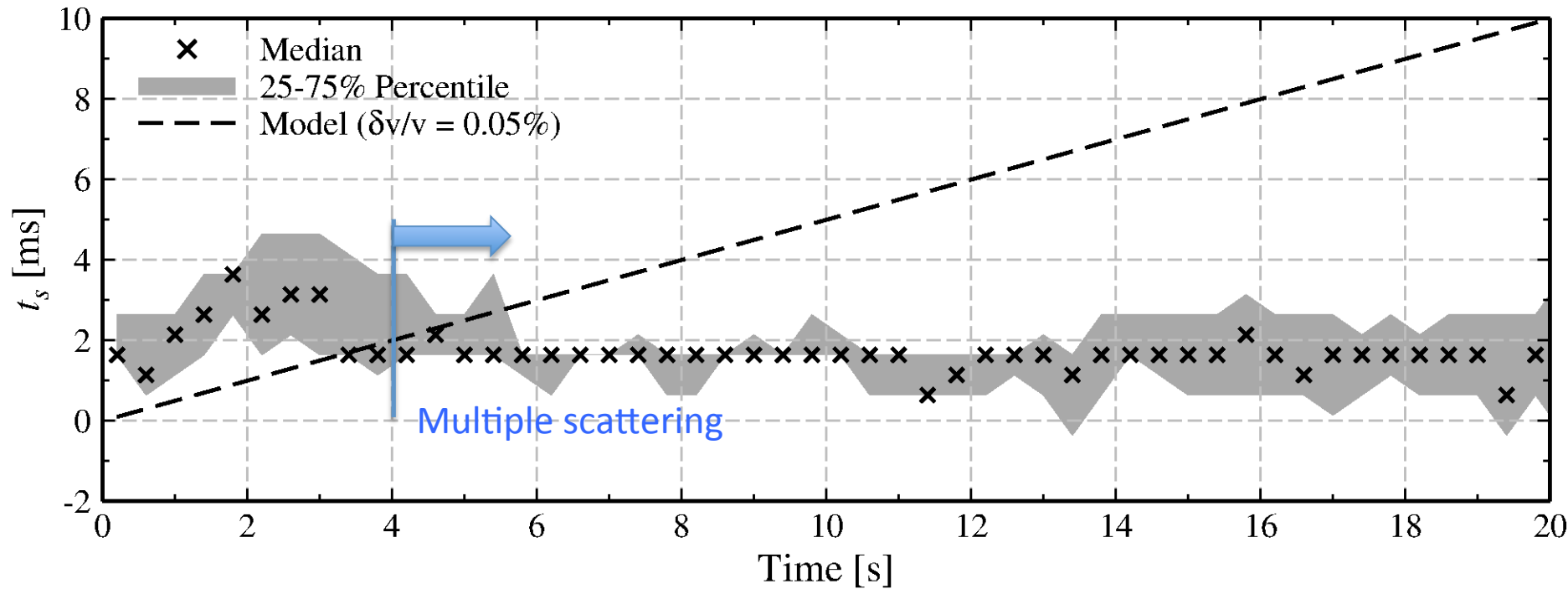


The seismic record (top panel), and estimates of δt for multiple time-windows (bottom panel). The first three points are computed from the early part of the waveform and are not sufficiently sensitive to detect the velocity change.

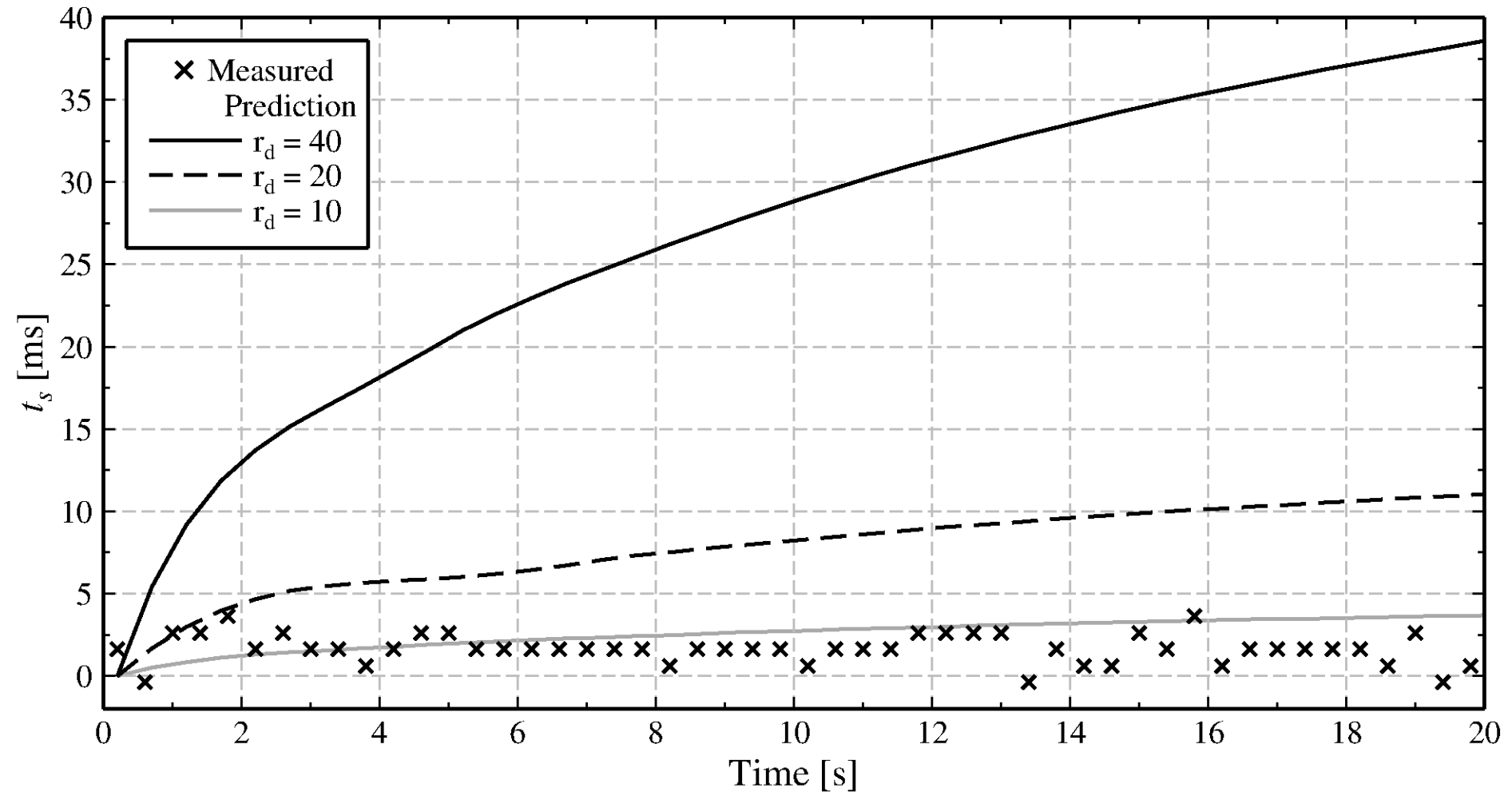
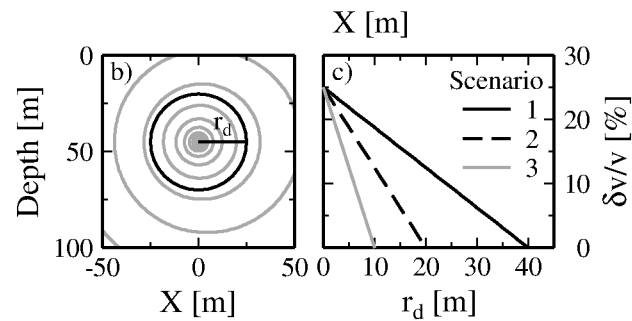
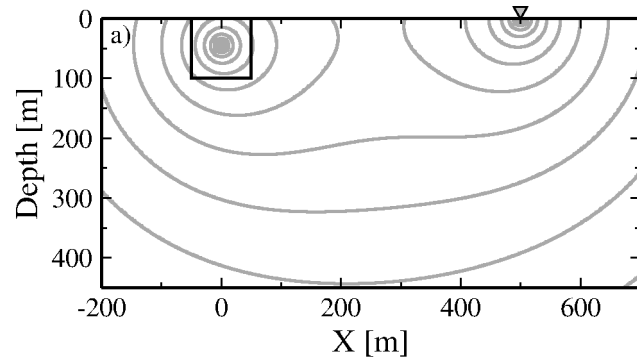
Analysis at L4-05 ($\Delta=500\text{m}$)



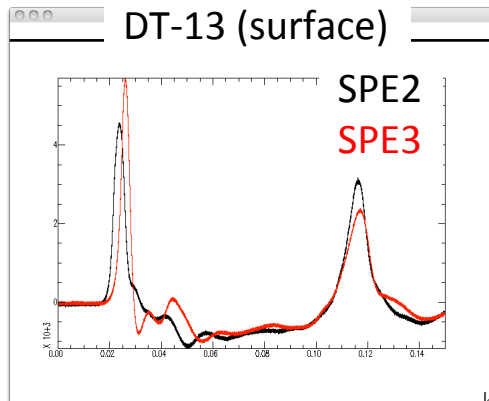
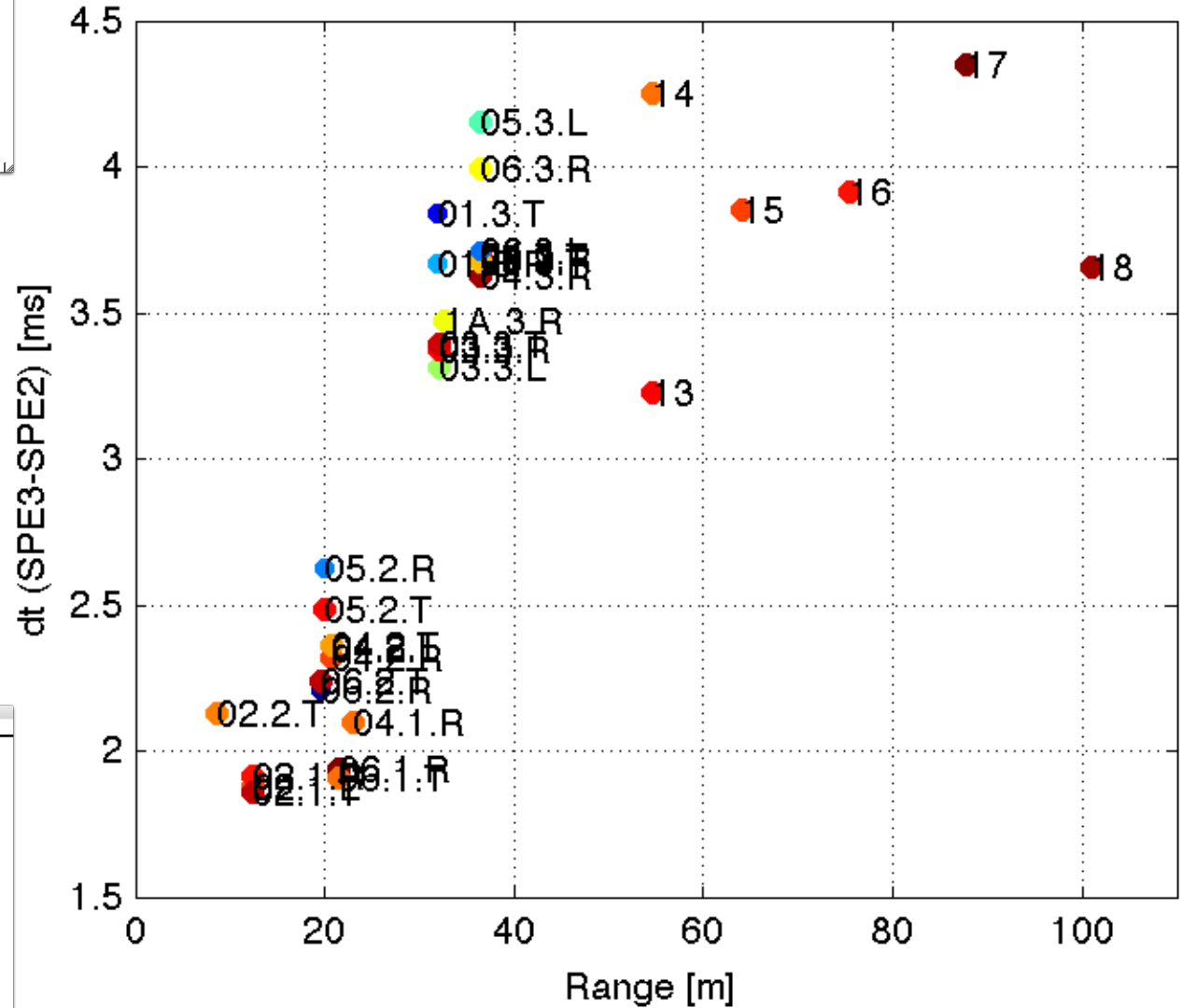
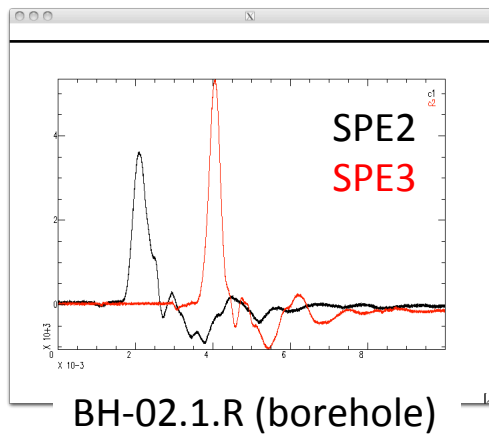
Analysis at all stations



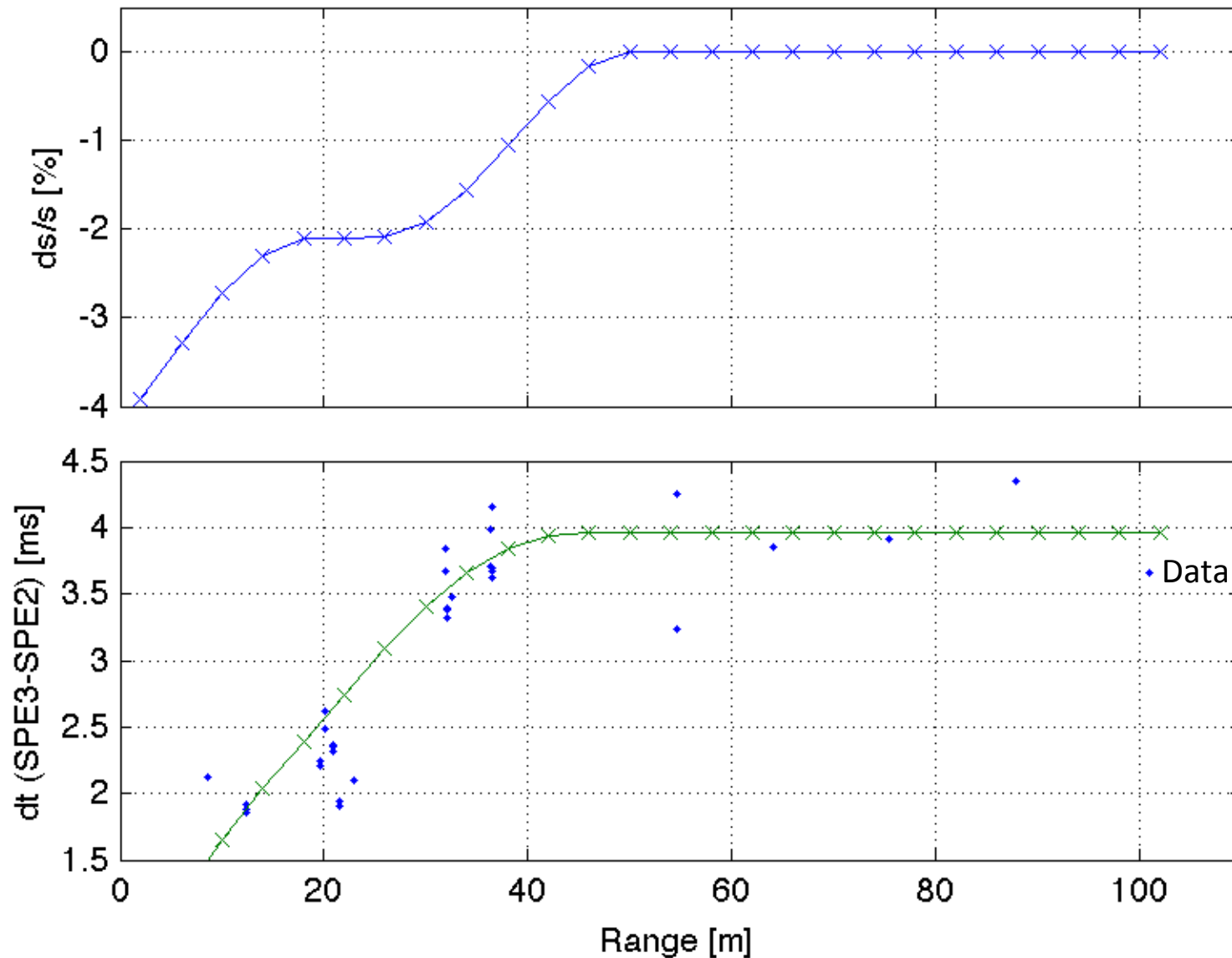
Prediction



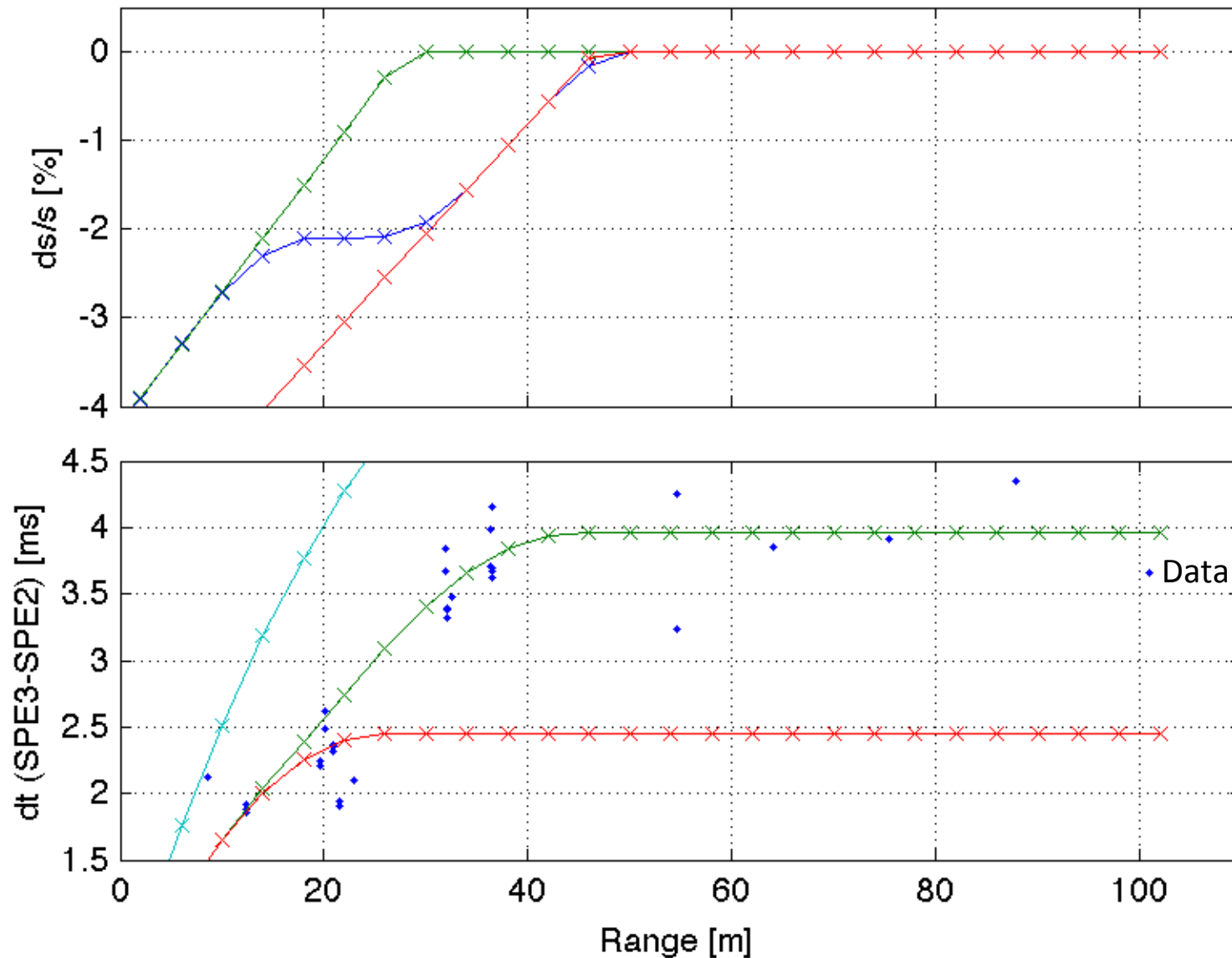
Near-field data



1D inversion for slowness perturbation



1D inversion for slowness perturbation



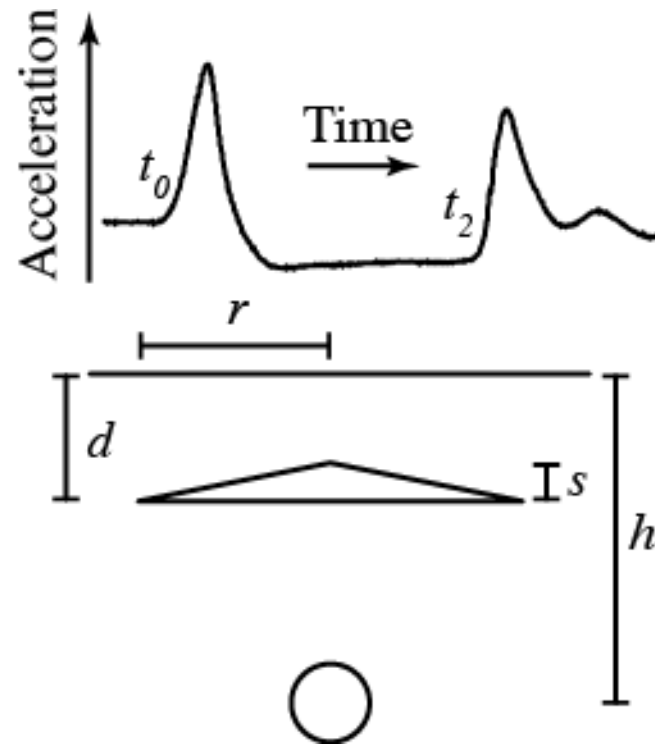
To Do

- Note SPE W and h
- Note historical parameters

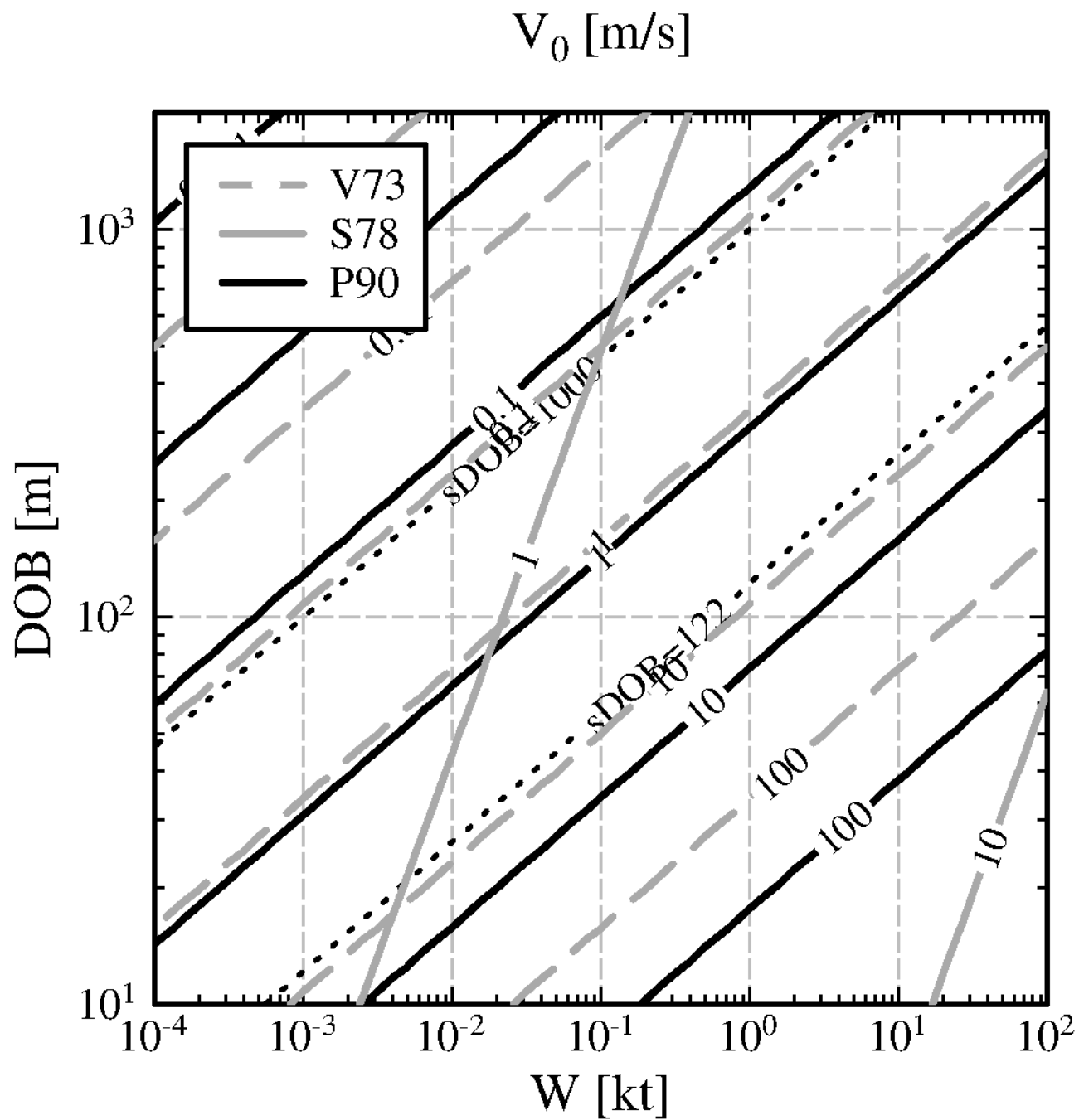
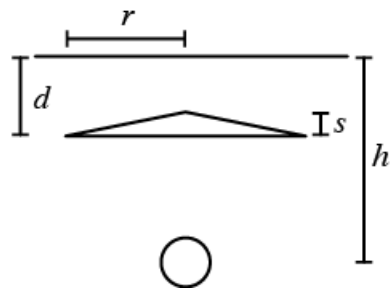
A Spall Model Comparison with Insights from the Source Physics Experiment

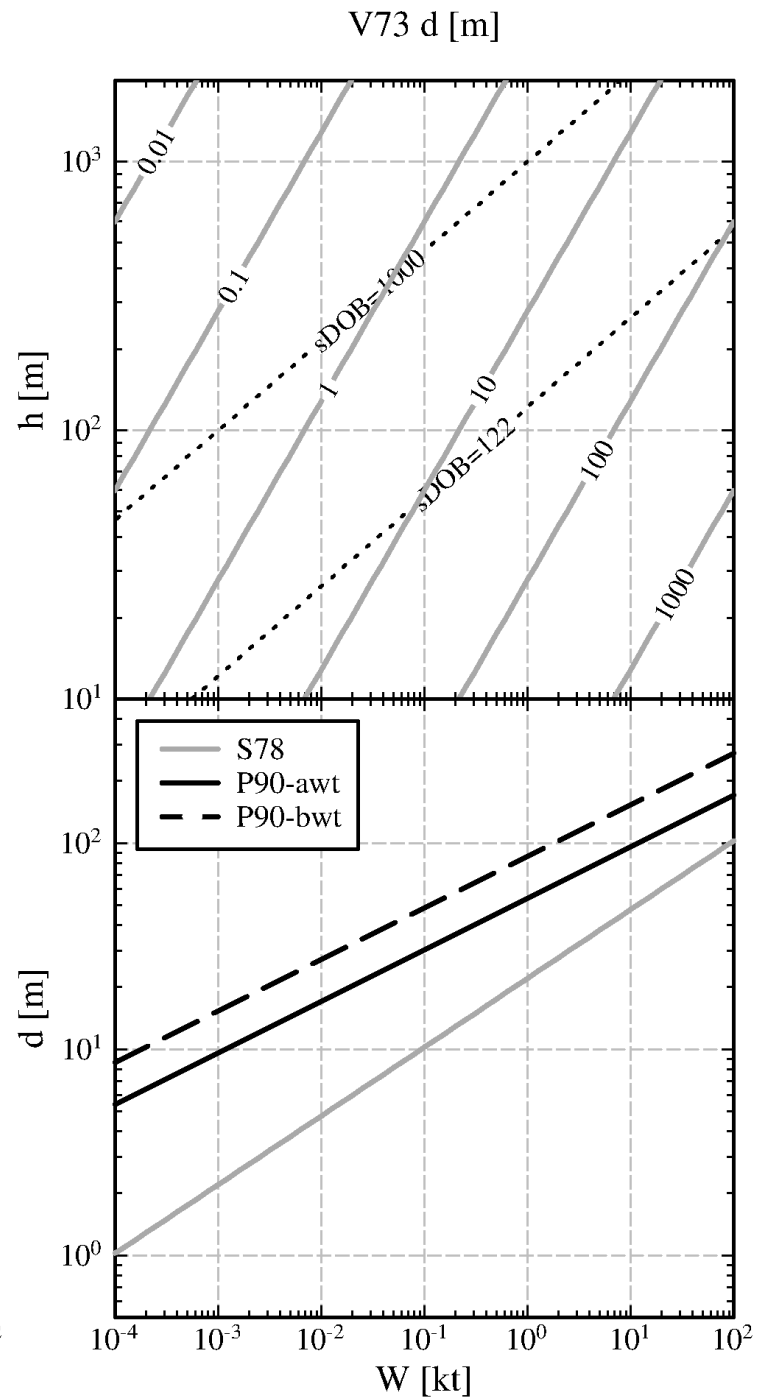
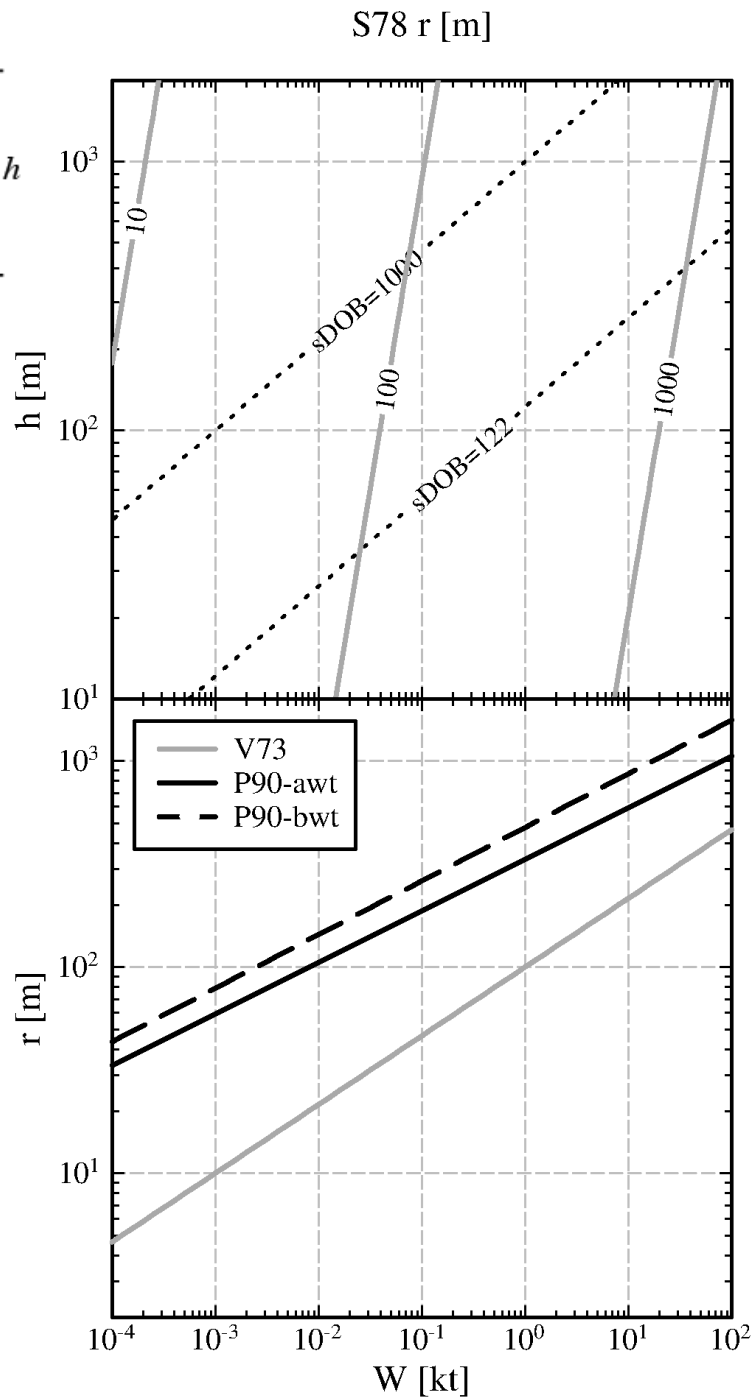
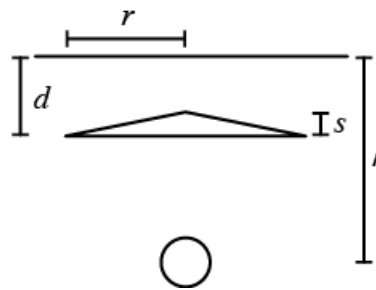
Sean Ford and Rob Mellors, LLNL

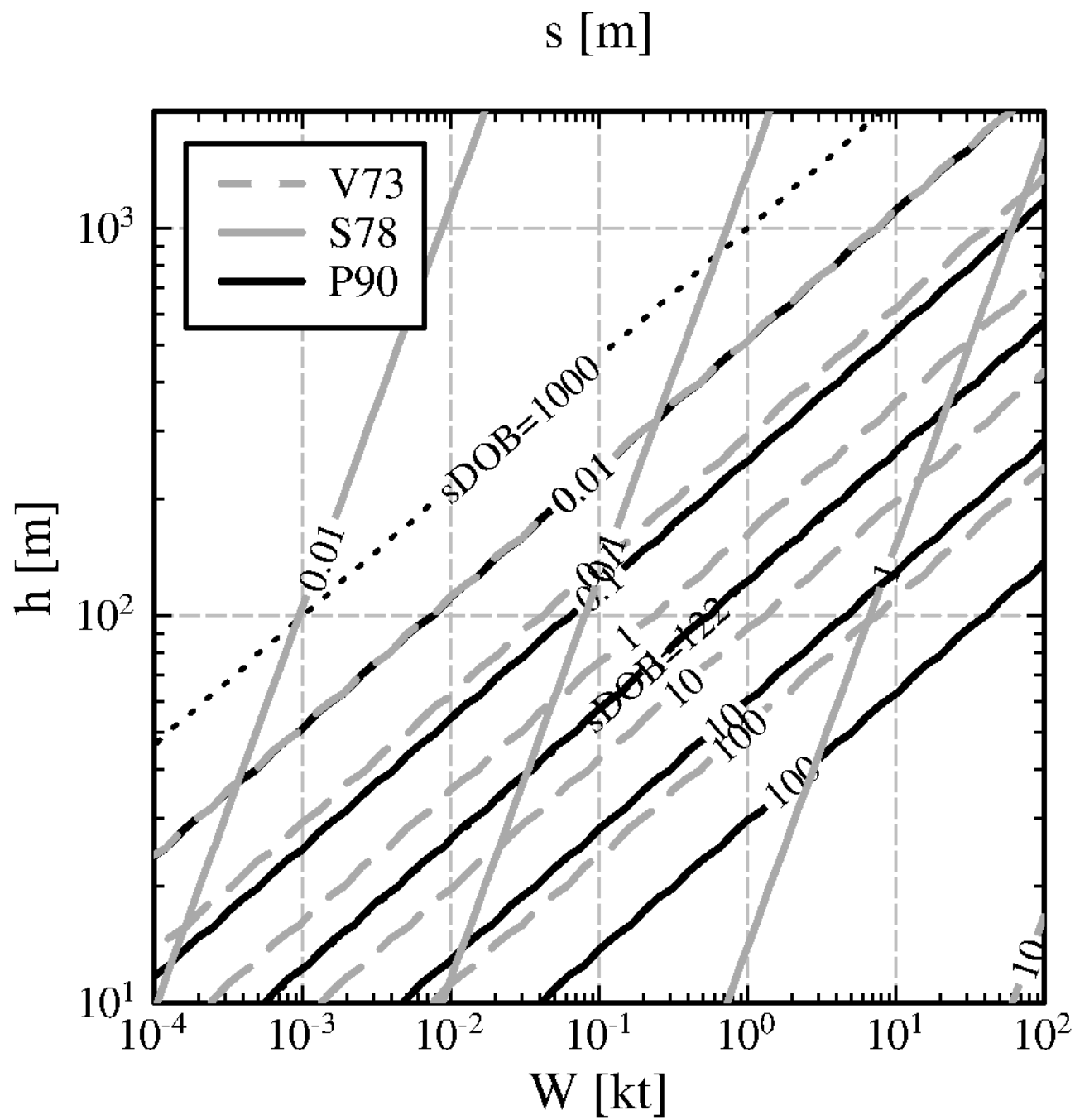
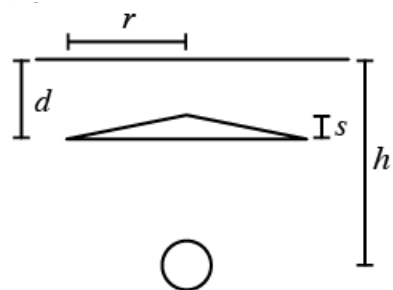
Spall model



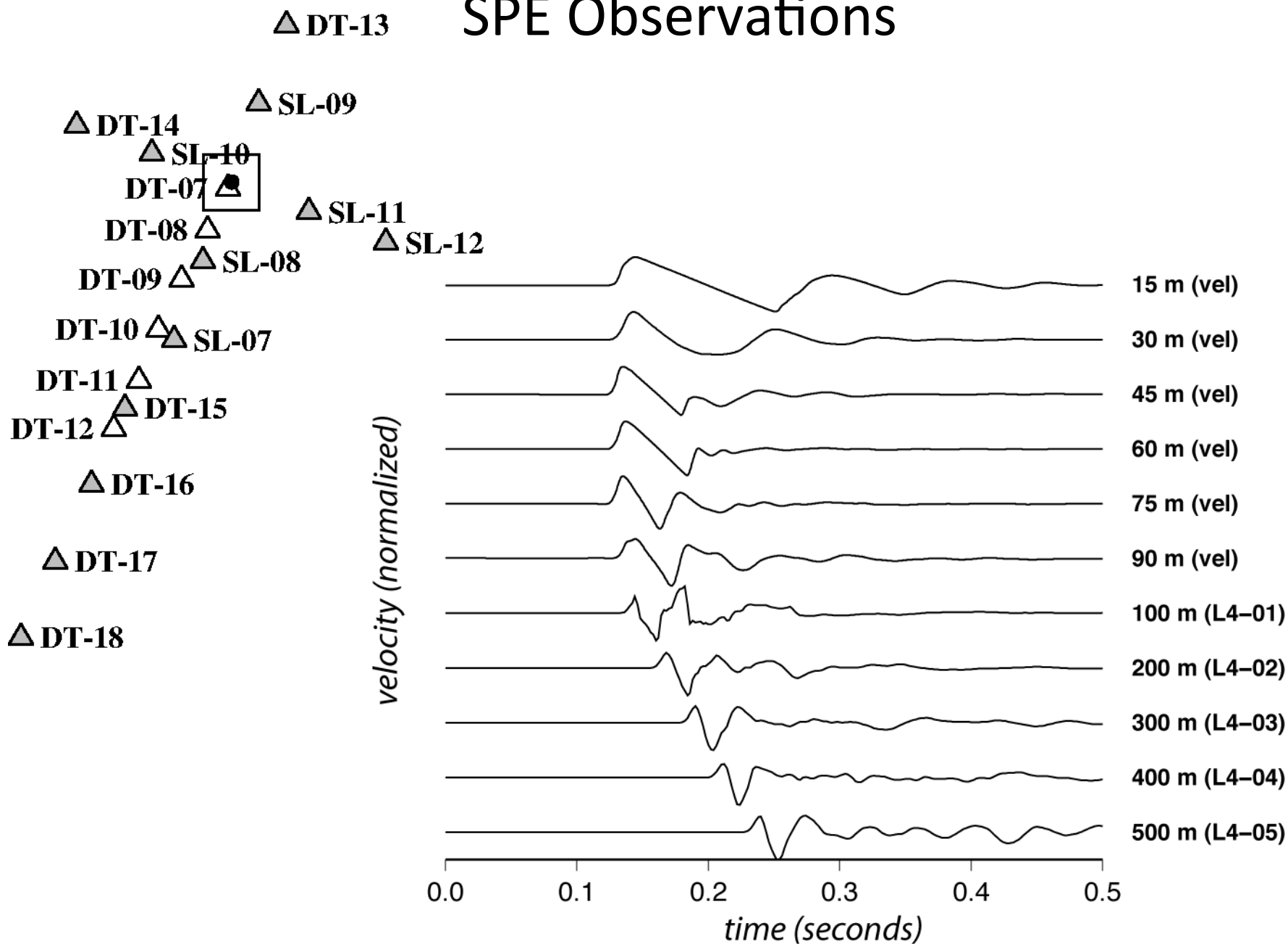
Study	Depth (d) [m]	Lateral extent (r) [m]	Maximum gap (s) [m]	Peak velocity (v) [m/s]
Viecelli (1973)	$10^{3.44} W^{2/3} h^{-1}$	$100 W^{1/3}$	$10^{8.82} W^{4/3} h^{-4}$	from s via eq (2)
Sobel (1978)	$22 W^{1/3}$	$10^{2.84} W^{0.37} h^{-0.16}$	$10^{0.57} W^{0.52} h^{-0.50}$	from s via eq (2)
Patton (1990)*	$54 W^{0.25}$	$334 W^{0.25}$	from v via eq (1)	$10^4 W^{1.61/3} h^{-1.61}$







SPE Observations



Comparison with PILEDRIIVER

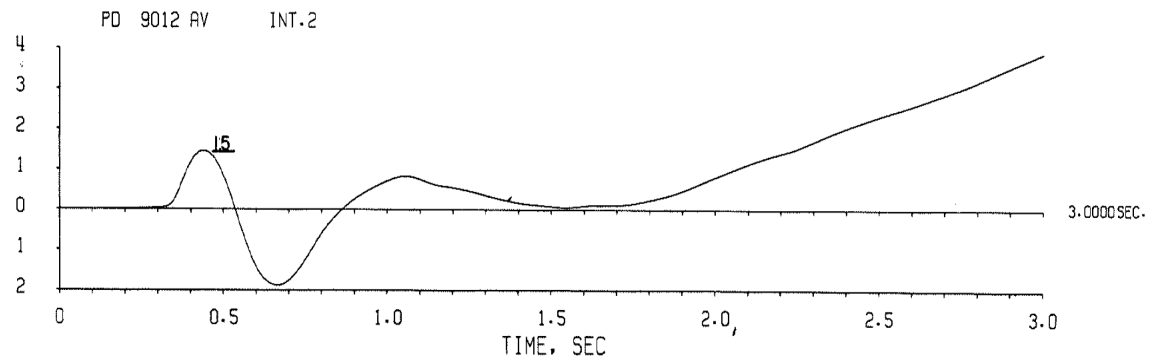
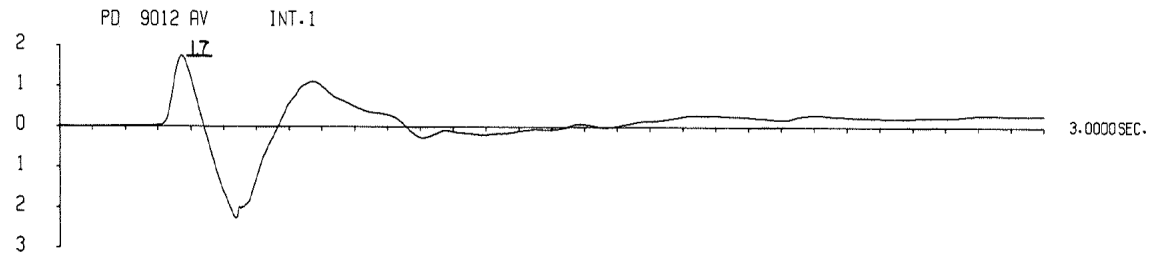
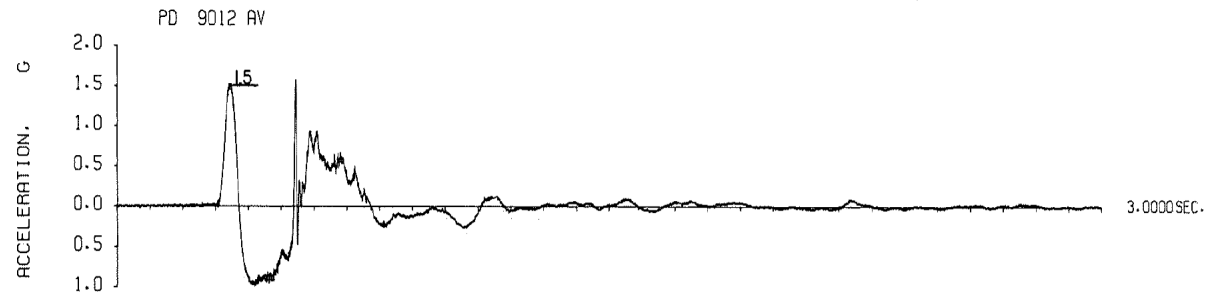
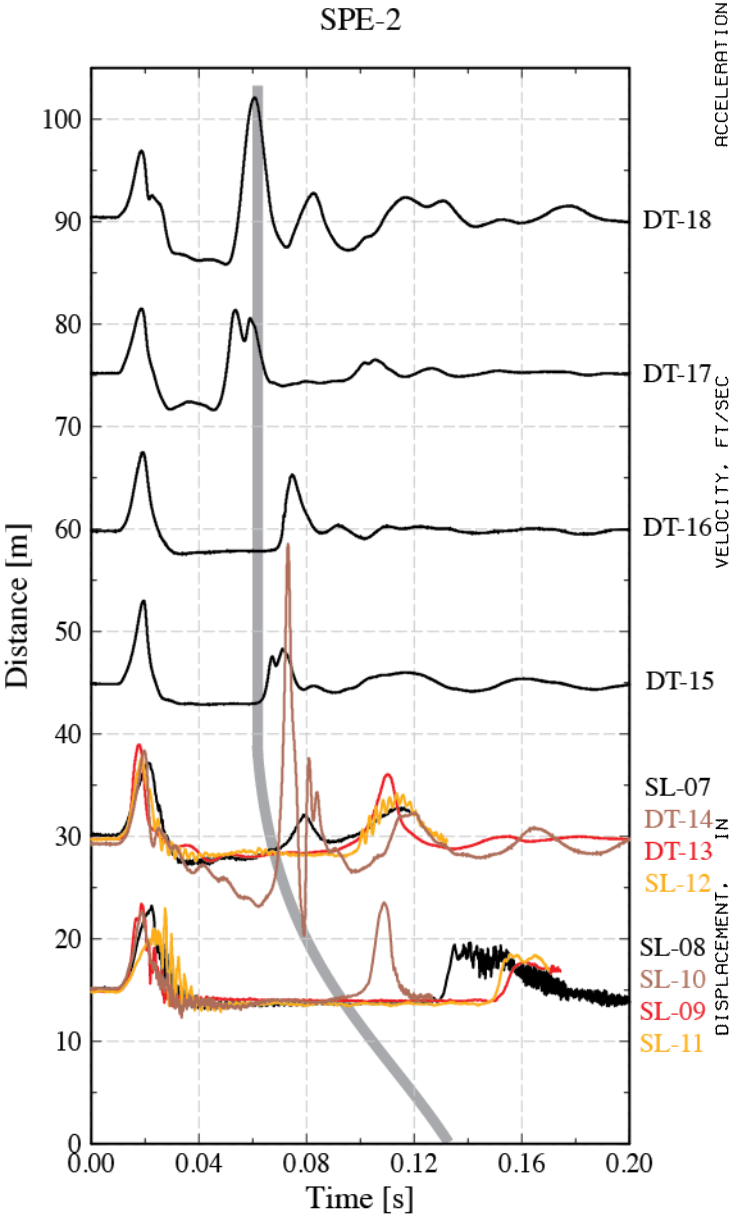


Figure 3.49 Vertical acceleration, velocity, and displacement, 4500 feet horizontal range, 1453 feet above the working point.

Comparison with PILEDRIIVER

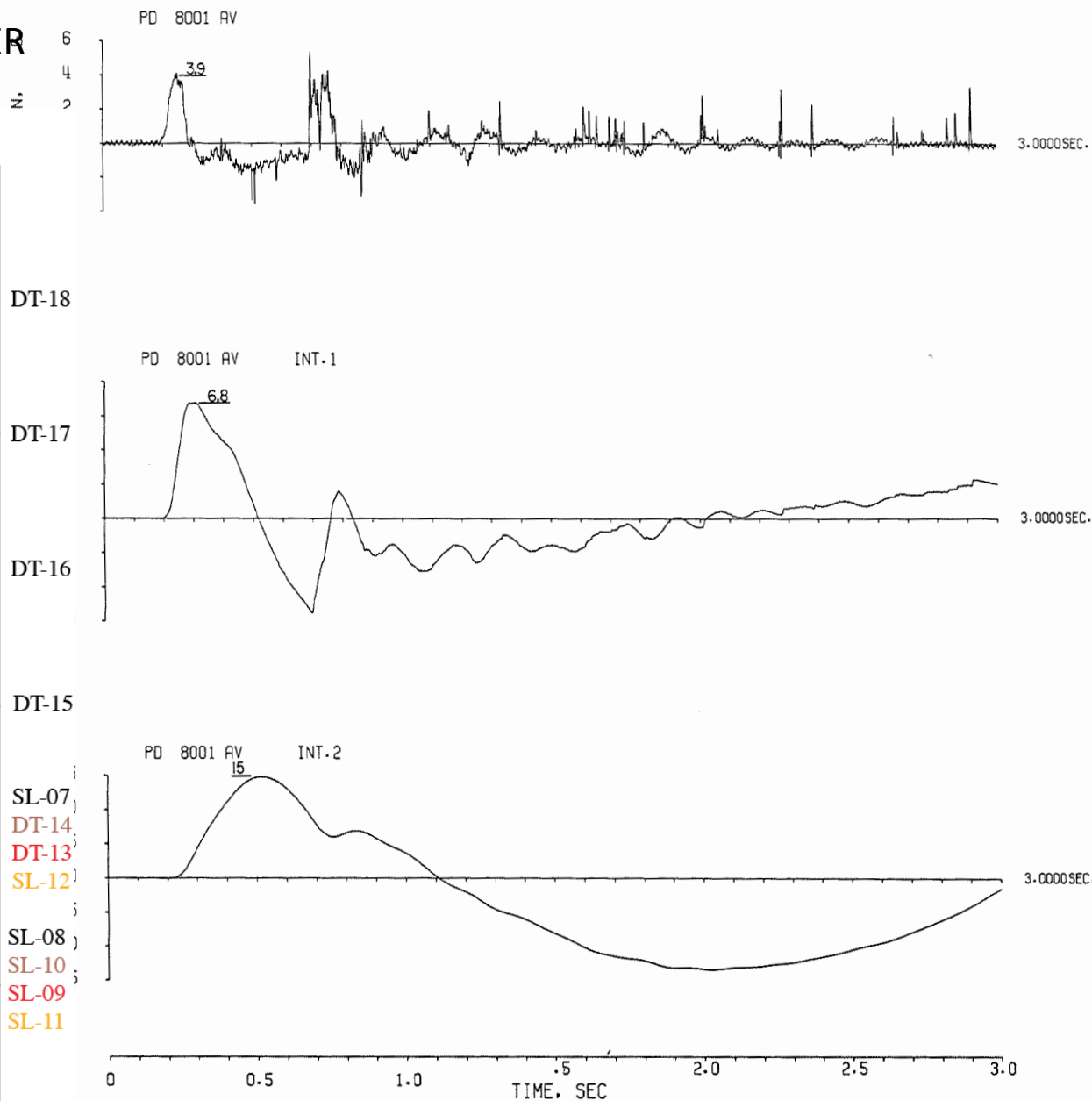
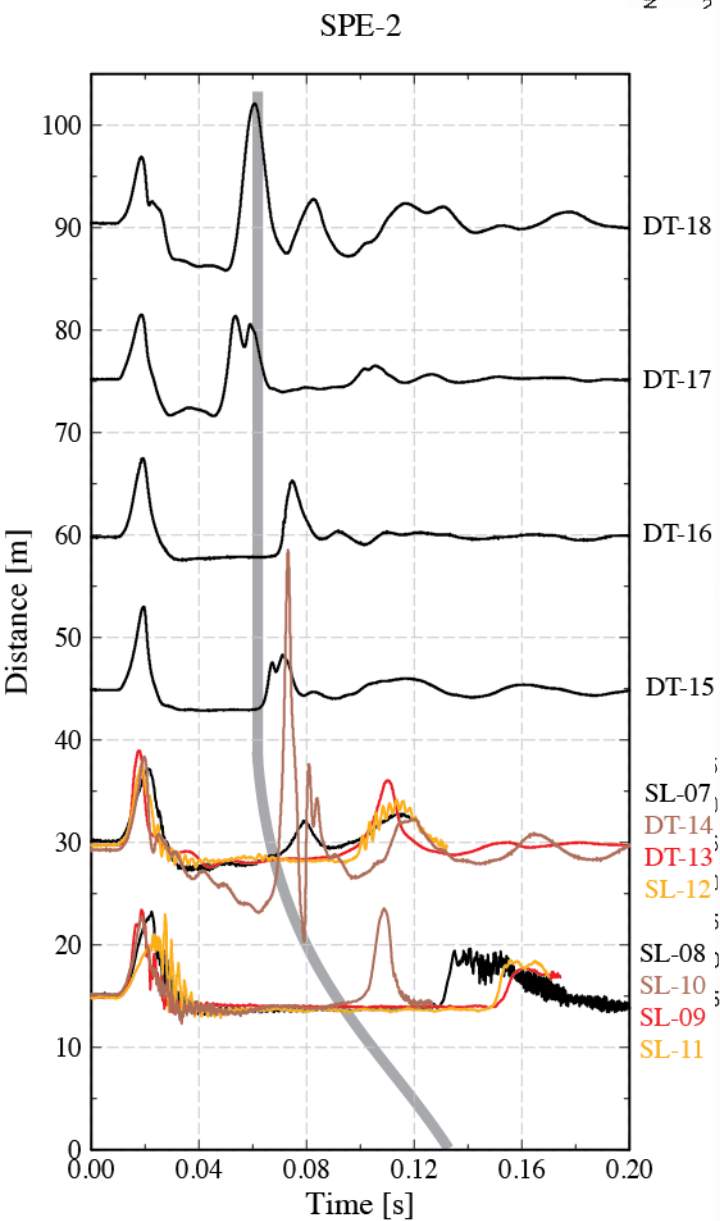


Figure 3.52 Vertical acceleration, velocity, and displacement, C van floor, front.

Comparison with PILEDRIVER

SPE-2

Distance [m]

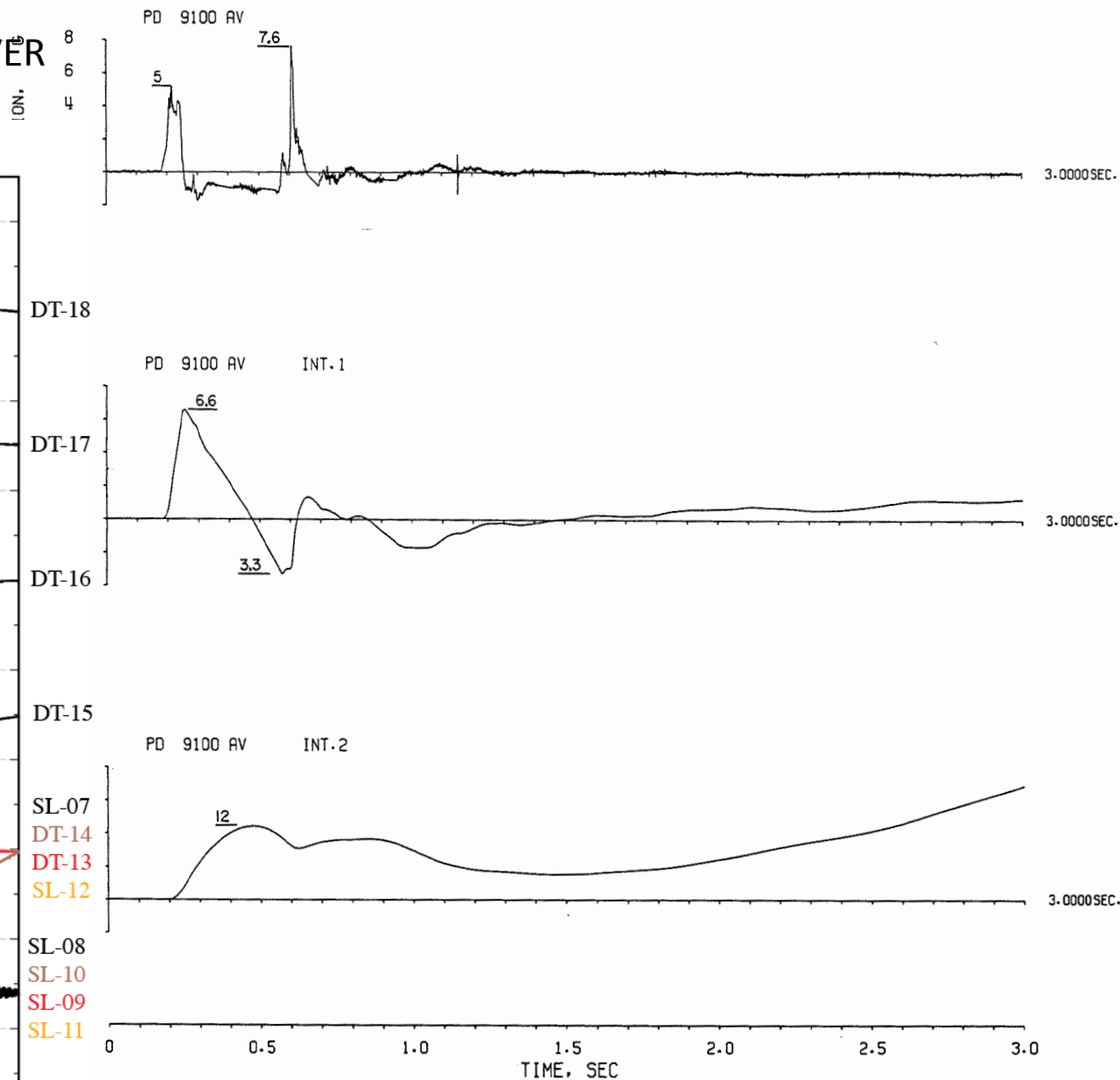
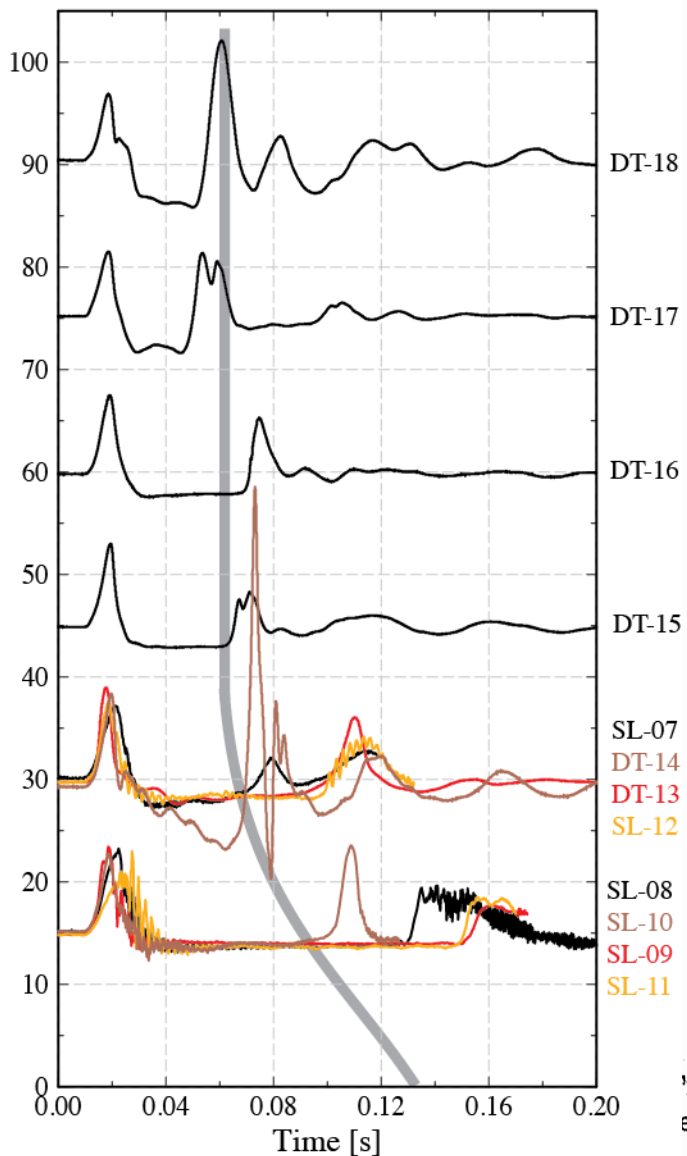


Figure 3.50 Vertical acceleration, velocity, and displacement in trailer park, 1850 feet horizontal range, 1409 feet above the working point.

Comparison with PILEDRIIVER

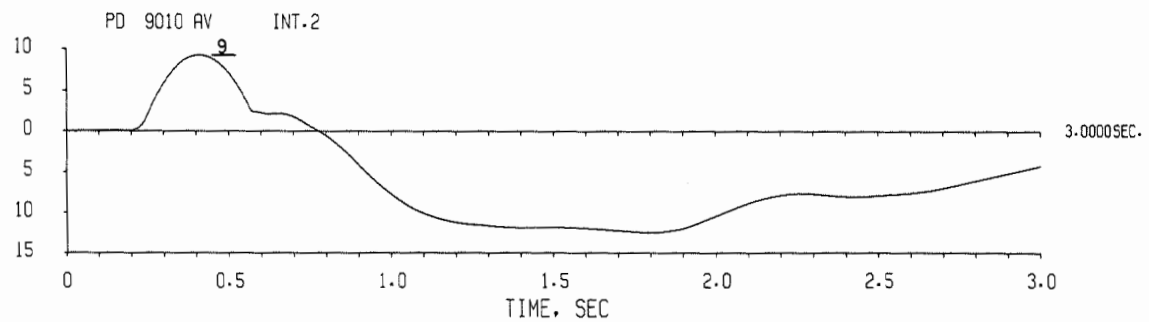
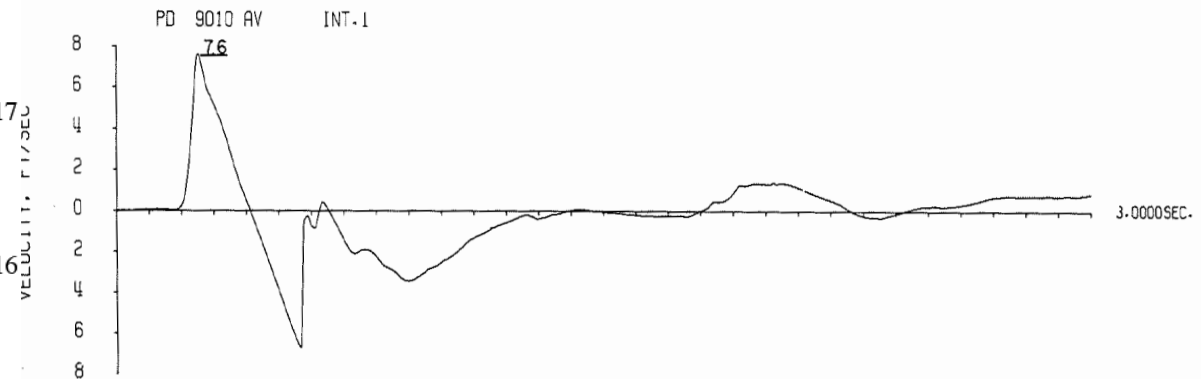
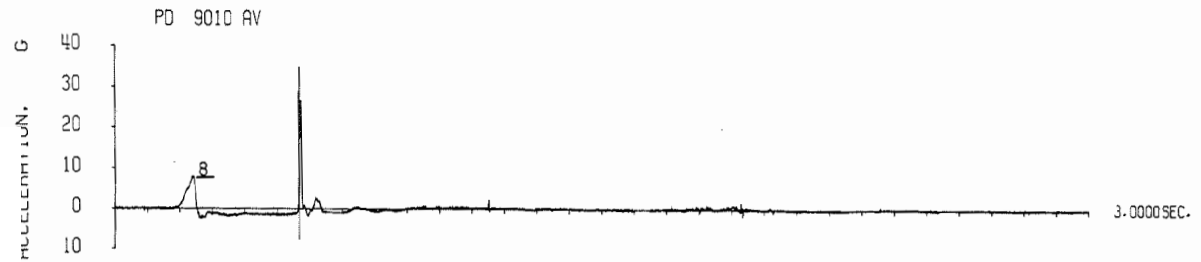
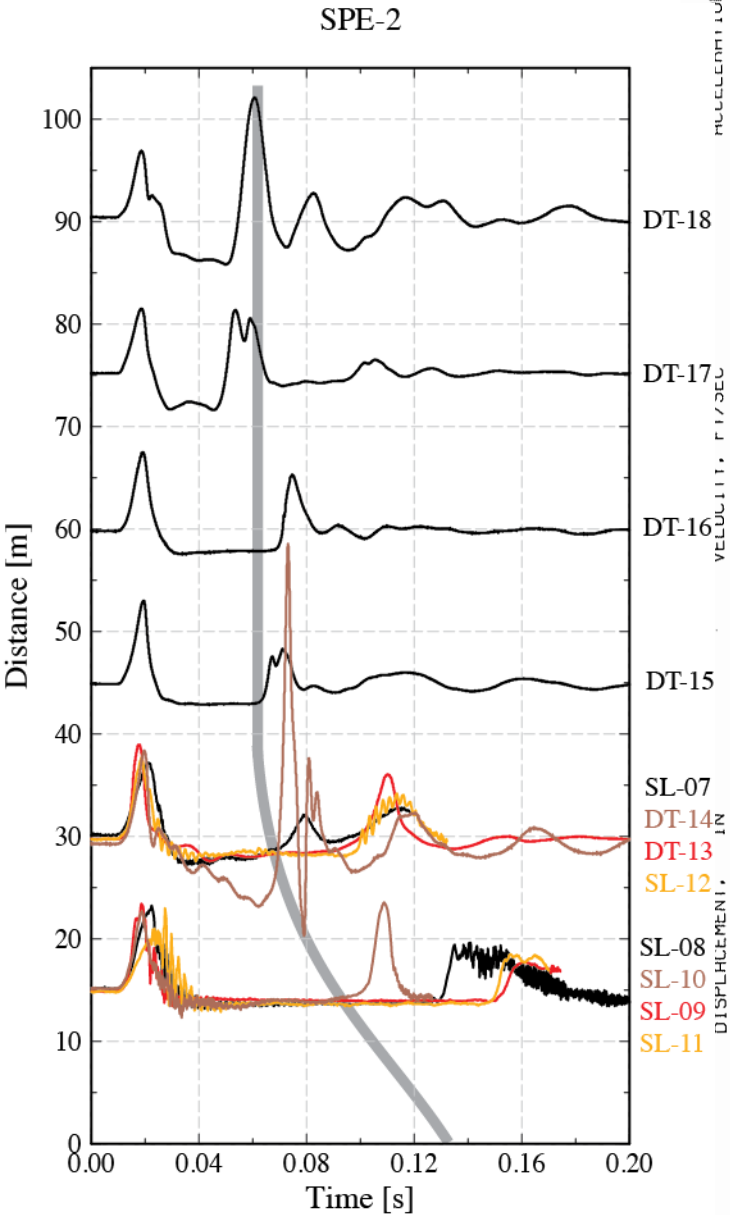
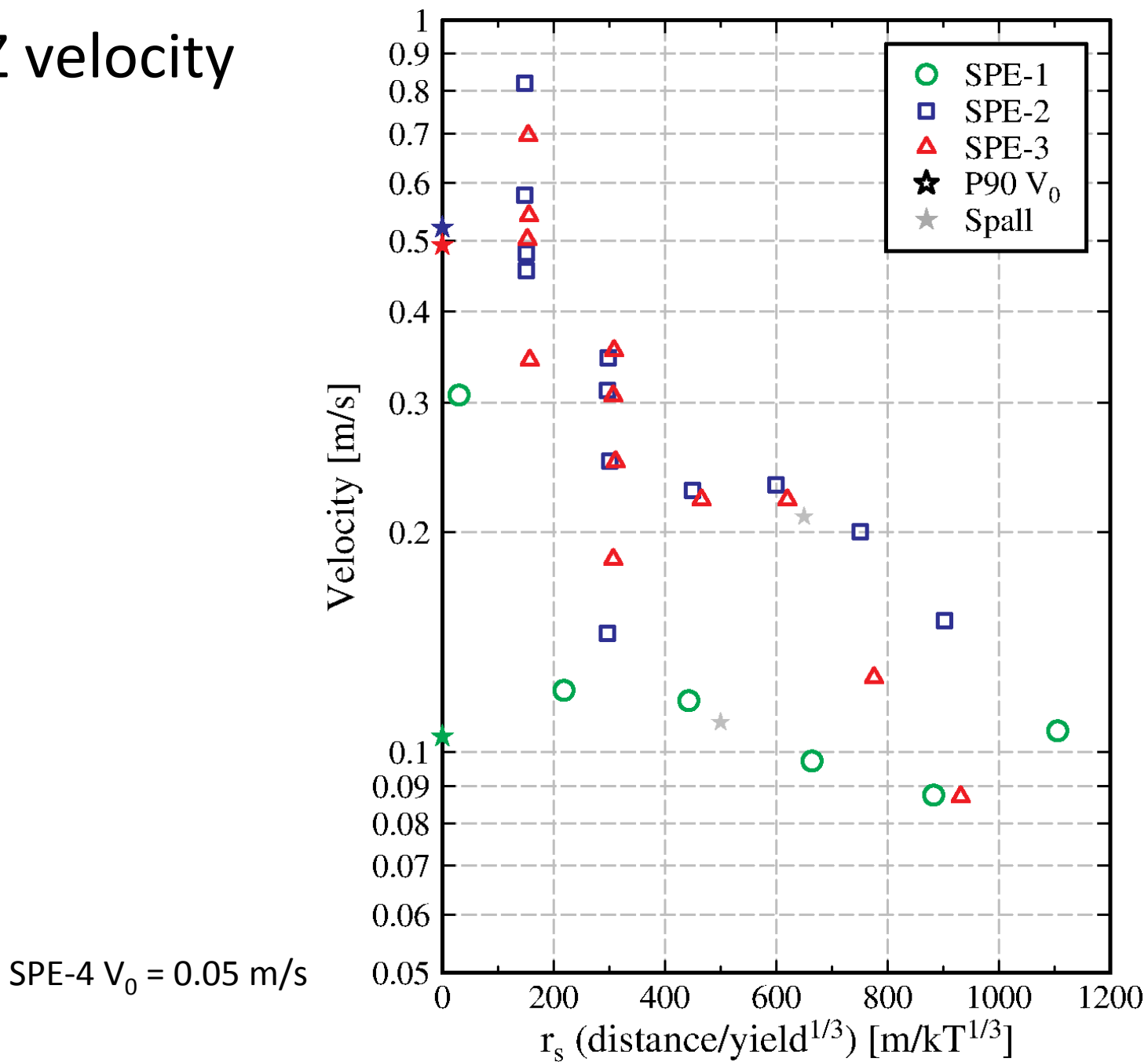
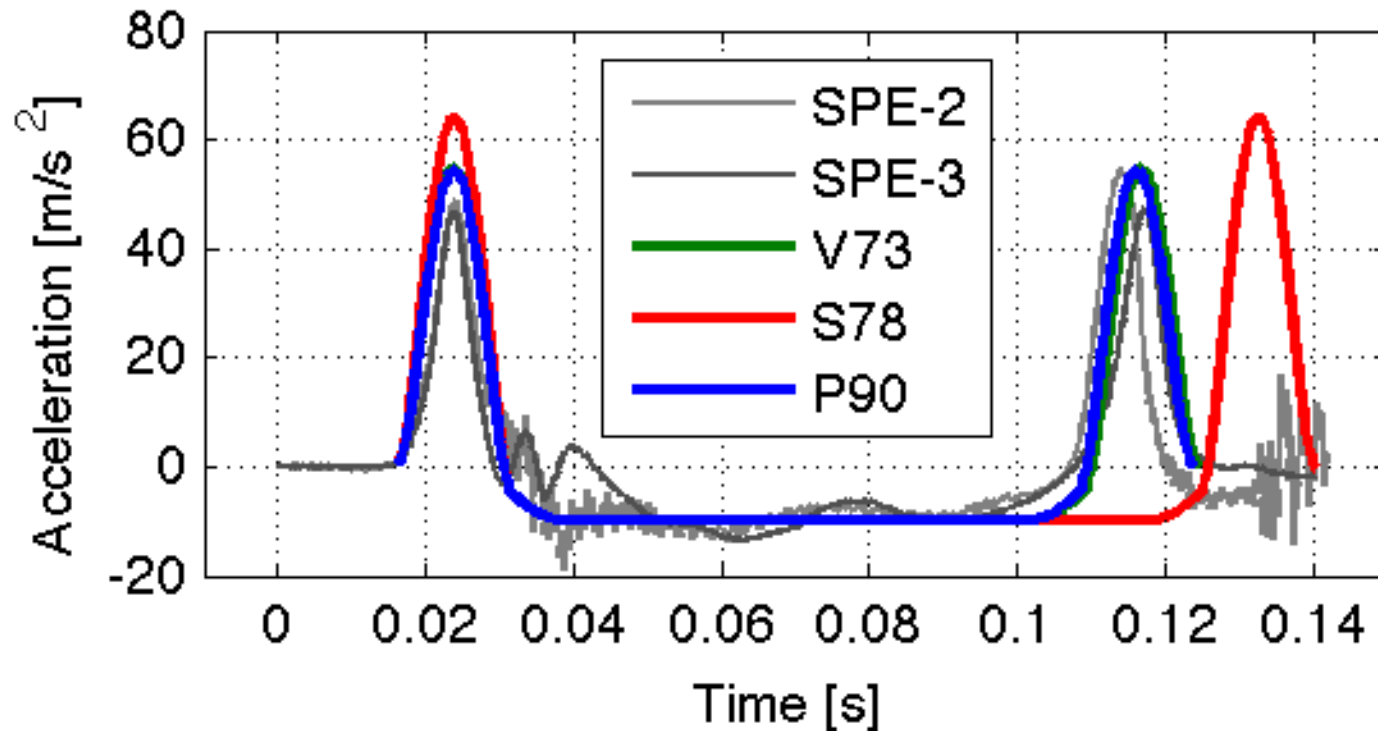


Figure 3.48 Vertical acceleration, velocity, and displacement, 2440 feet horizontal range, 1539 feet above the working point.

Peak GZ velocity

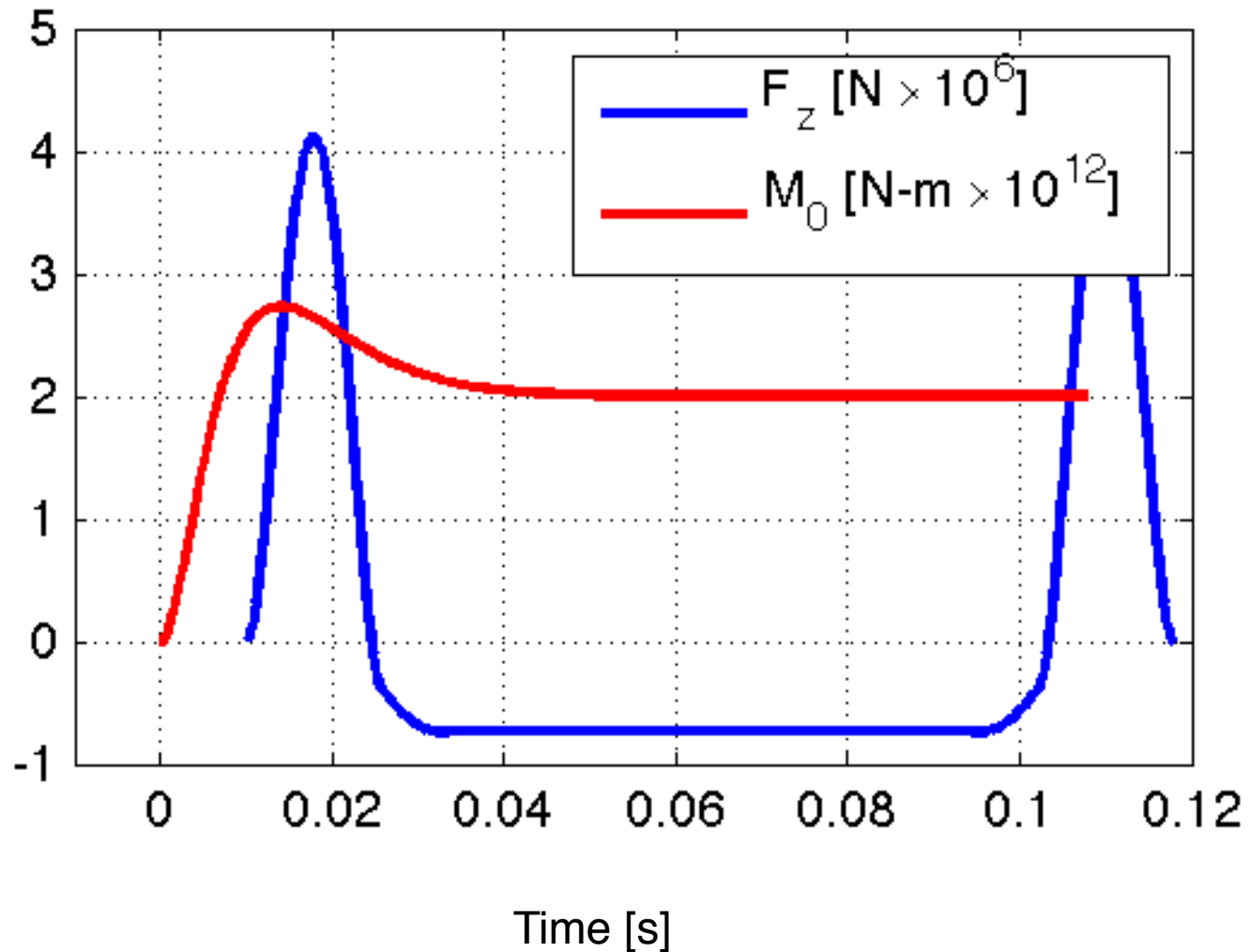


Model comparison with SL-10 (15m-NW)

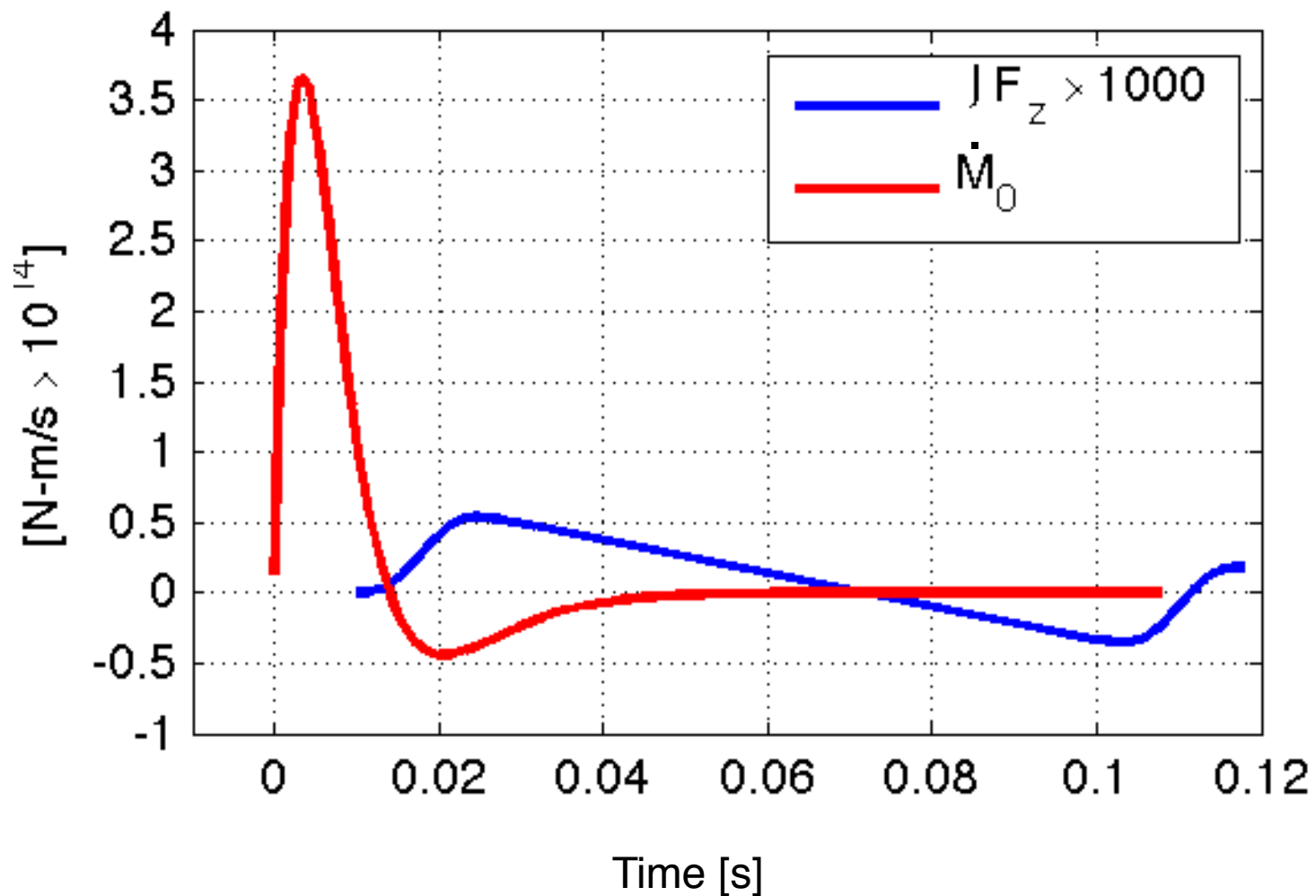


Model	Gap [cm]	Radius [m]	V_0 [cm/s]	Mass [T]	Dwell [ms]	Rise [ms]
V73	1.06	10.00	45.51	2.18	92.9	15.6
S78	1.45	28.72	53.24	24.62	108.7	15.6
P90	1.04	59.39	45.15	75.73	92.2	15.6

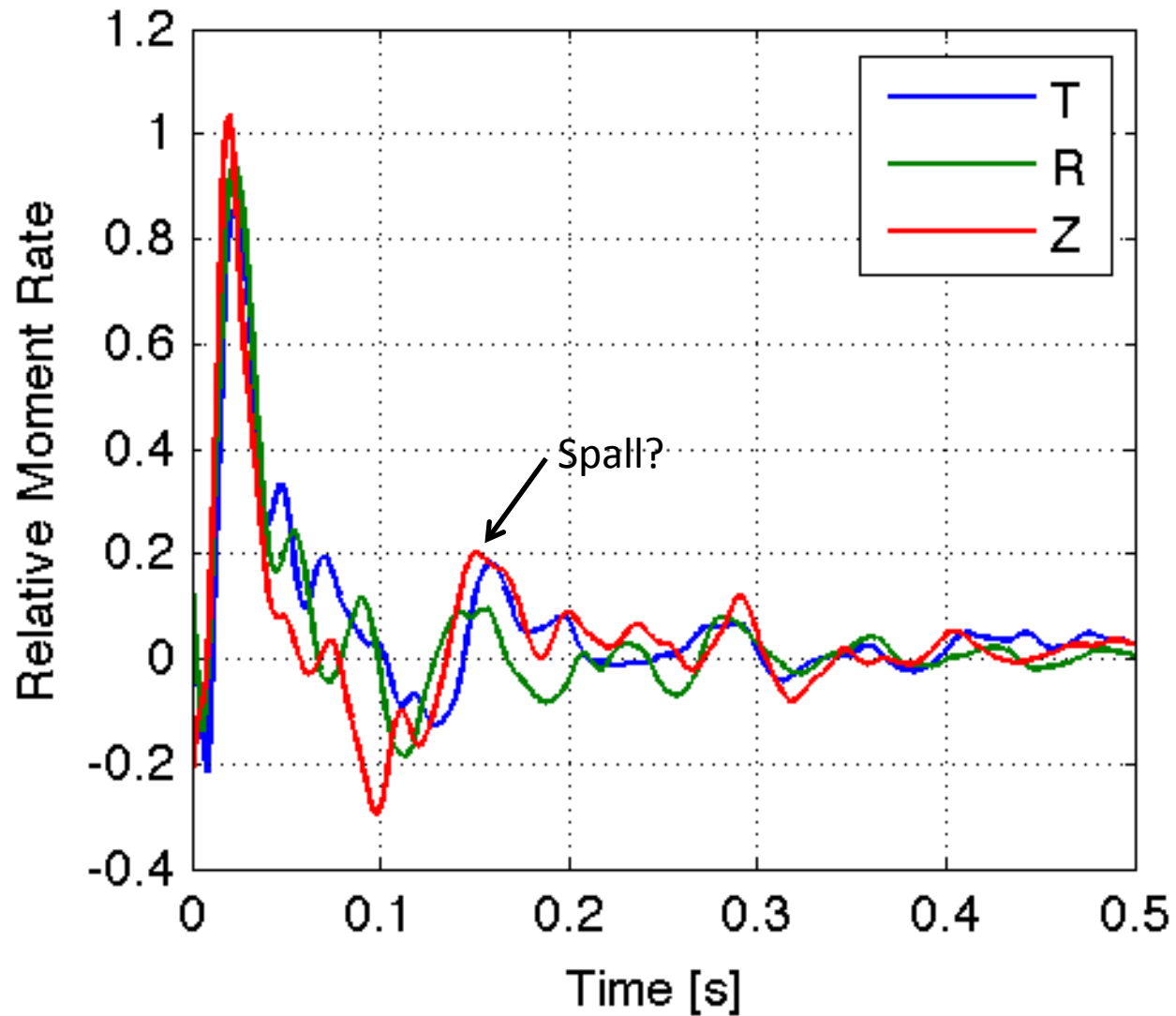
Explosion moment and spall force functions



Far-field displacement functions



SPE-2/-1 Deconvolution



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